

A visualization of the cosmic web, showing a complex network of dark matter filaments and voids. Numerous bright yellow and orange stars are scattered throughout the structure, with some clusters appearing denser than others. The overall color palette is dark, with shades of grey and black for the filaments, and bright yellow and orange for the stars.

Vegyí elemek, óriás fekete lyukak és egy darabokra hullott műhold

Csillagoktól az atomokig

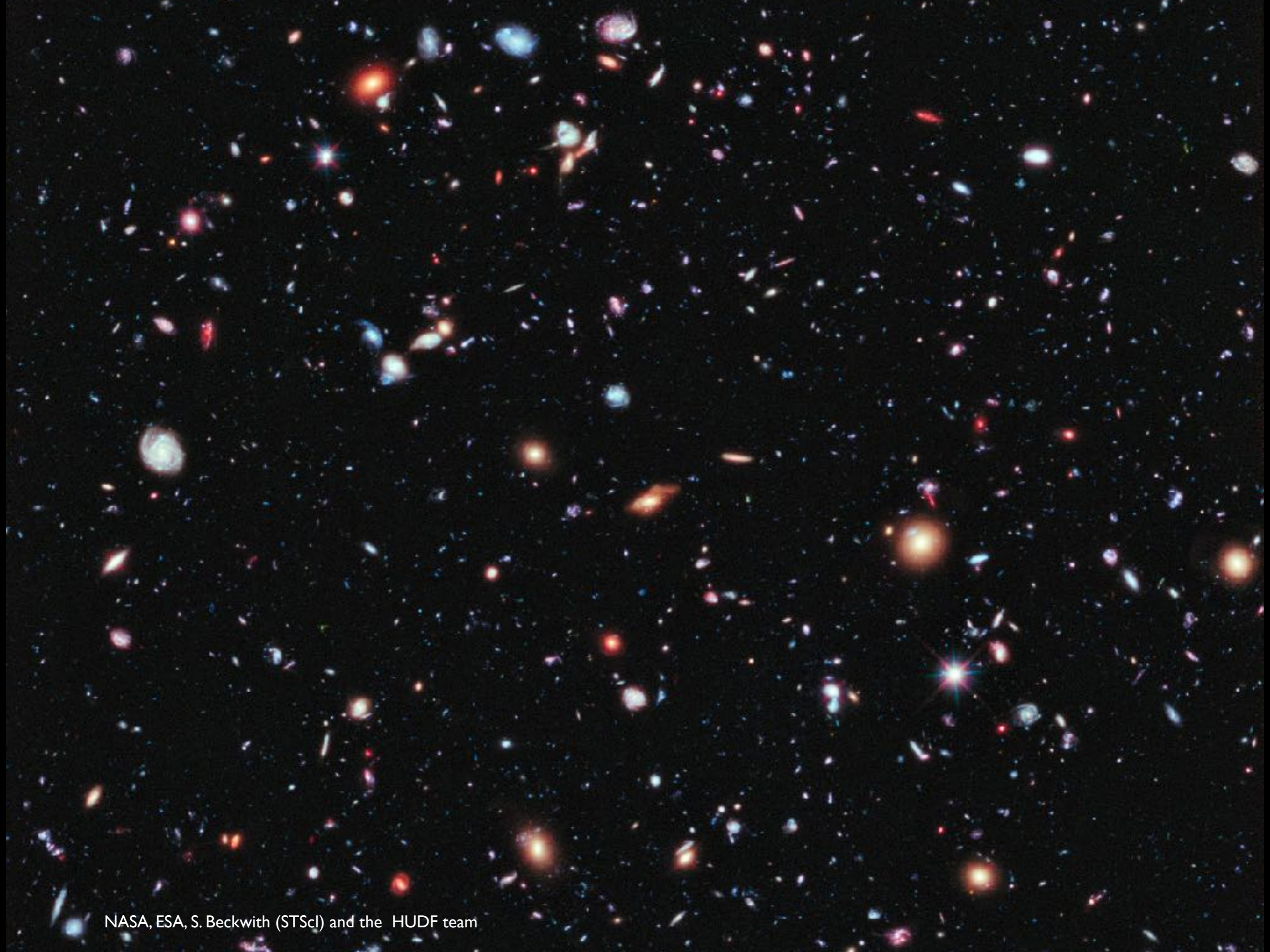
Werner Norbert



Jason Lau

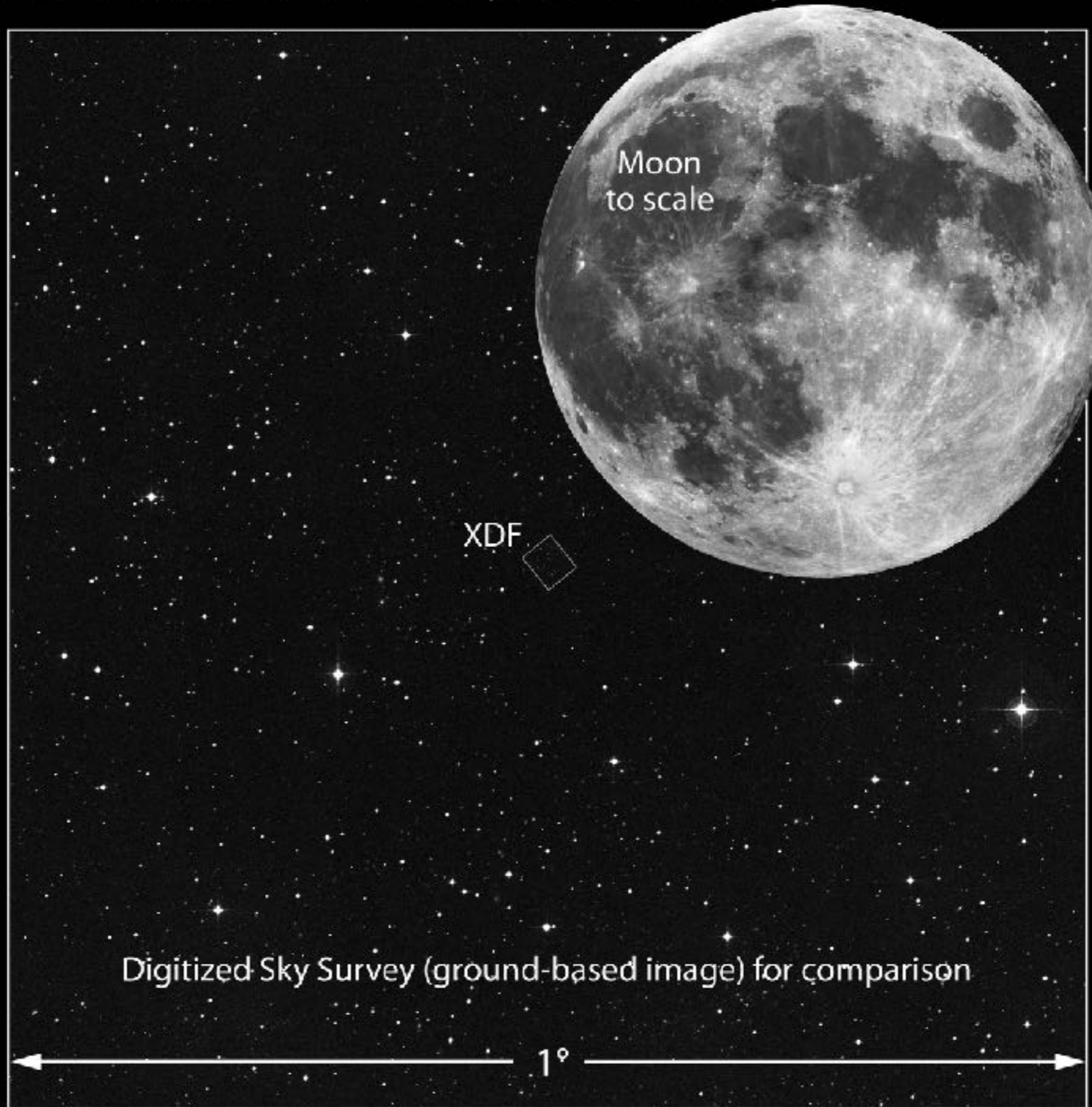


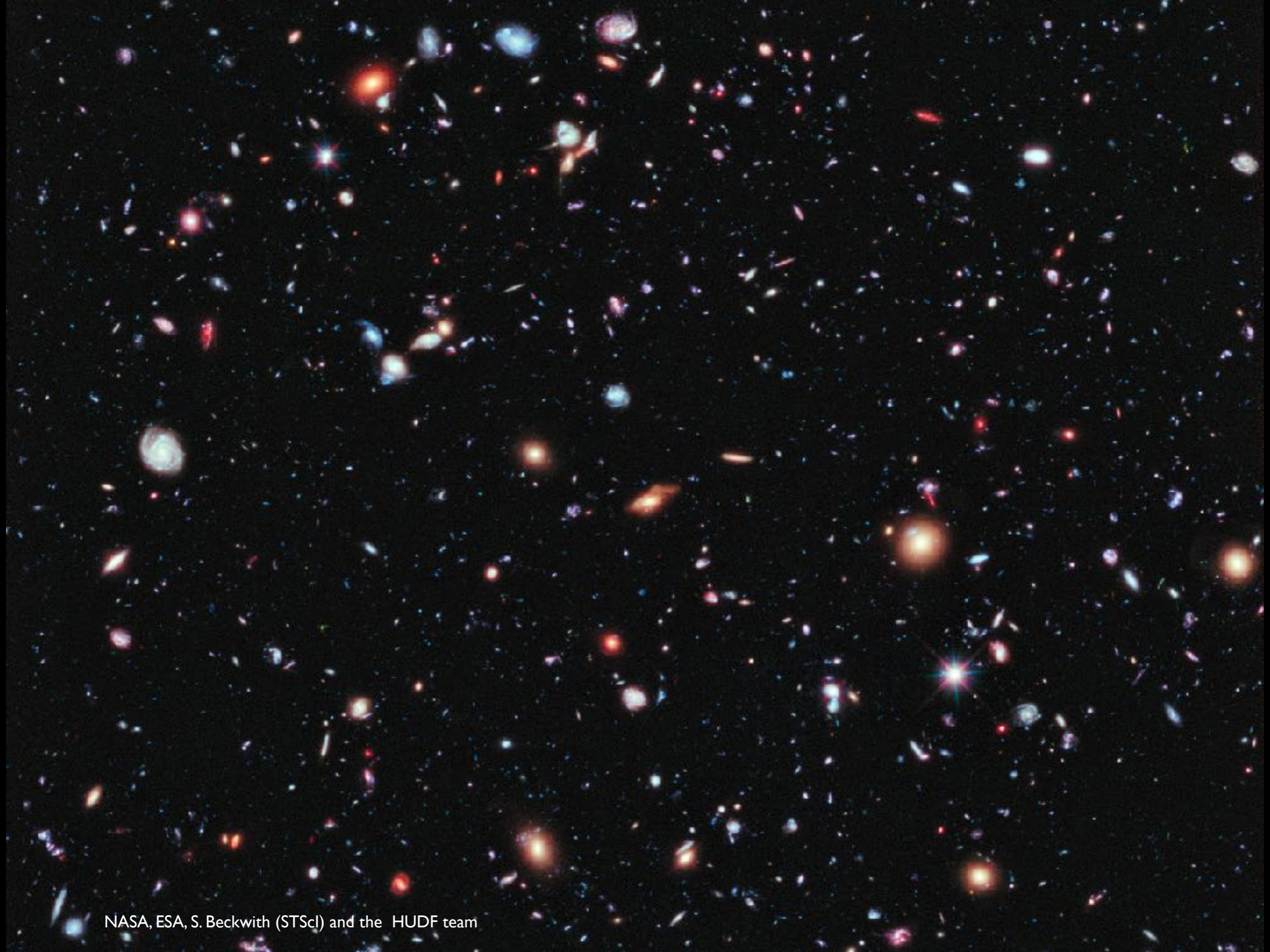
Credit: Nick Risinger



NASA, ESA, S. Beckwith (STScI) and the HUDF team

Size of Hubble eXtreme Deep Field on the Sky

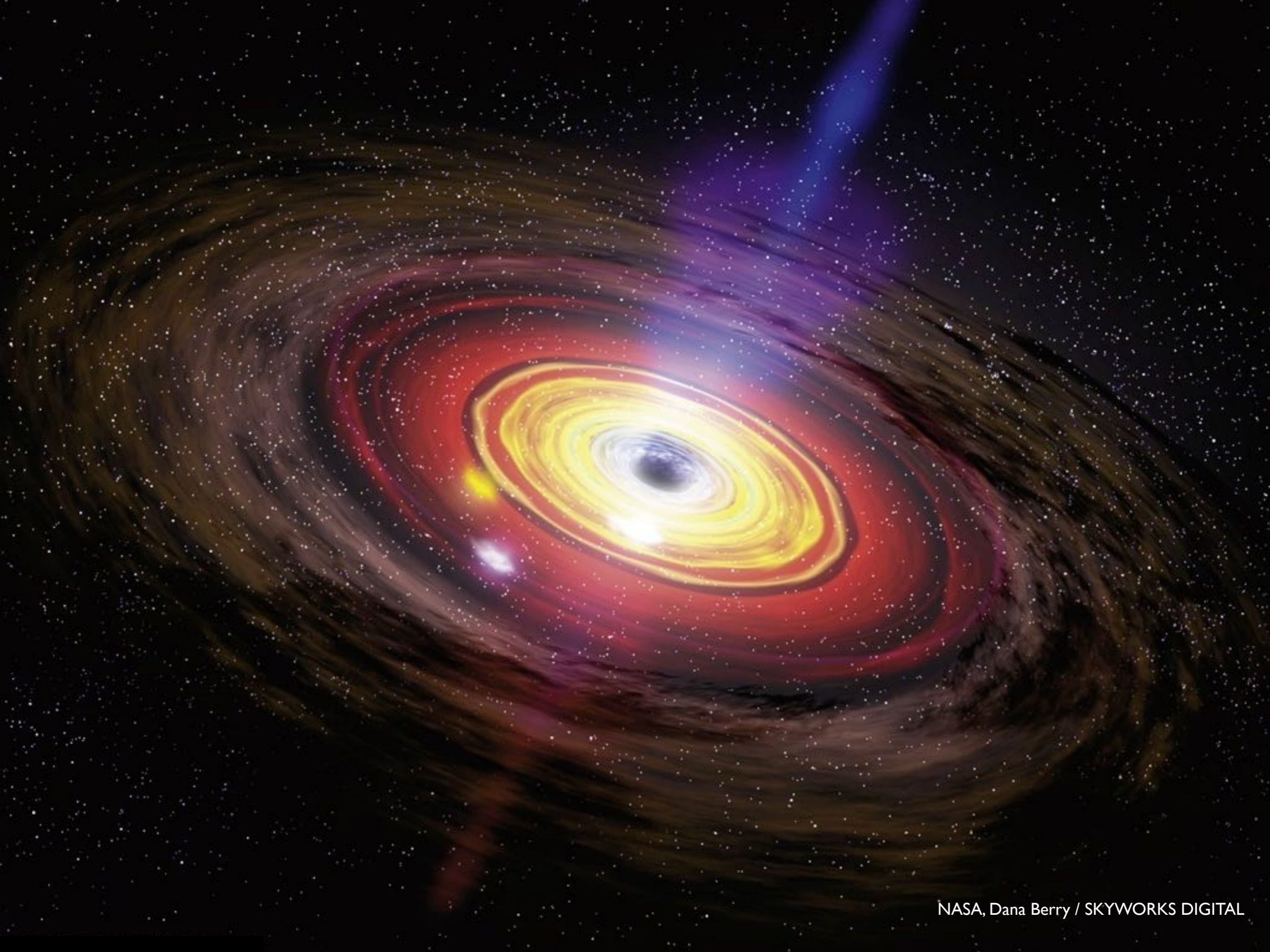




NASA, ESA, S. Beckwith (STScI) and the HUDF team



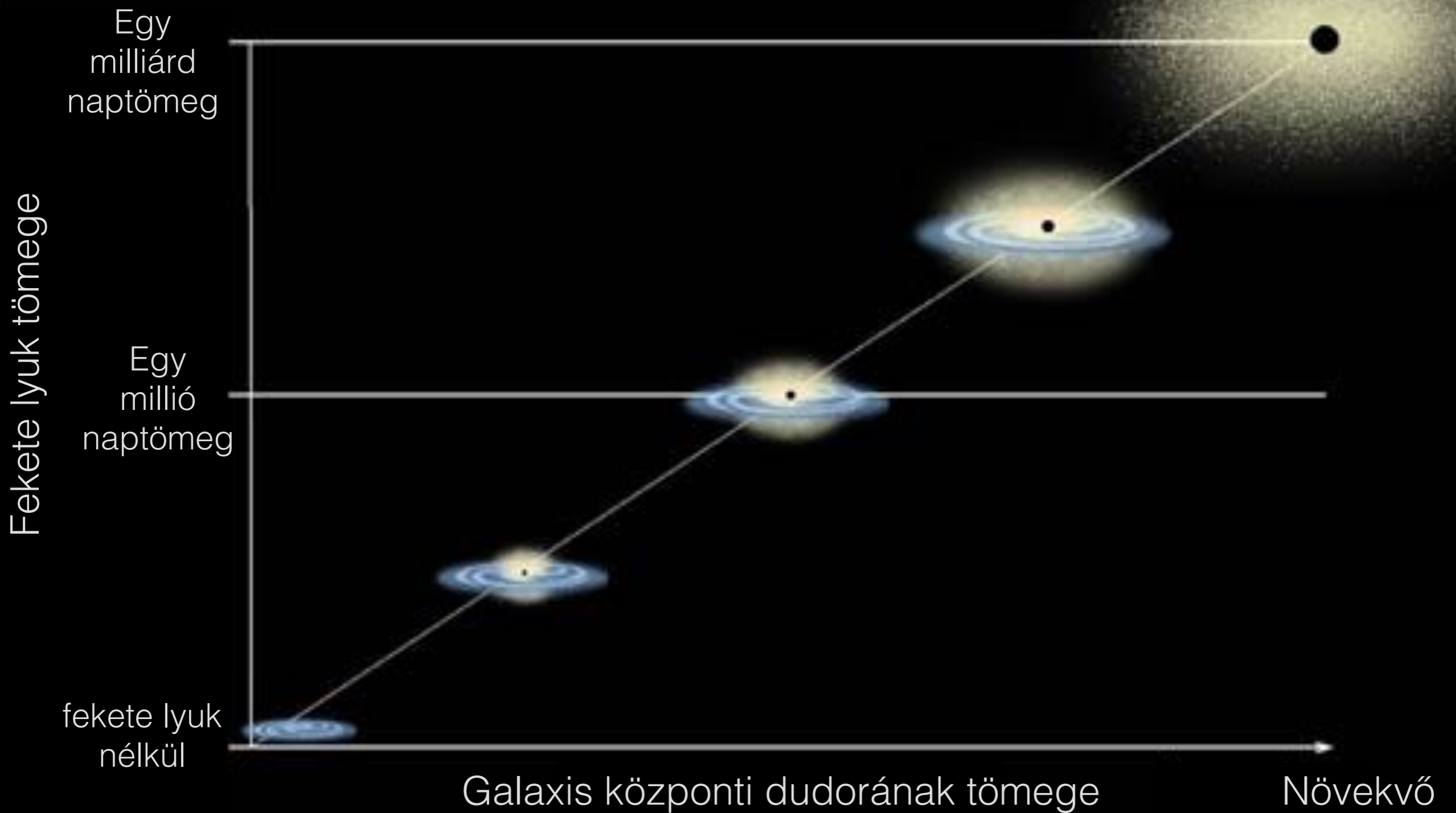


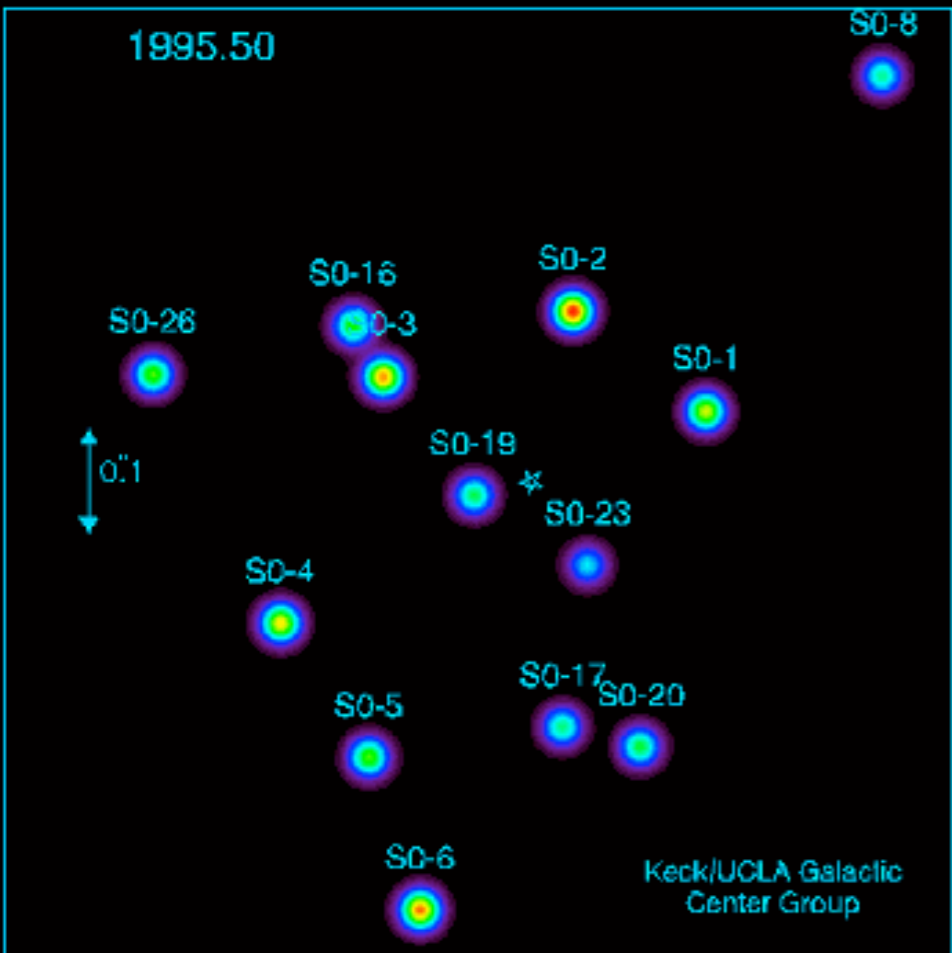


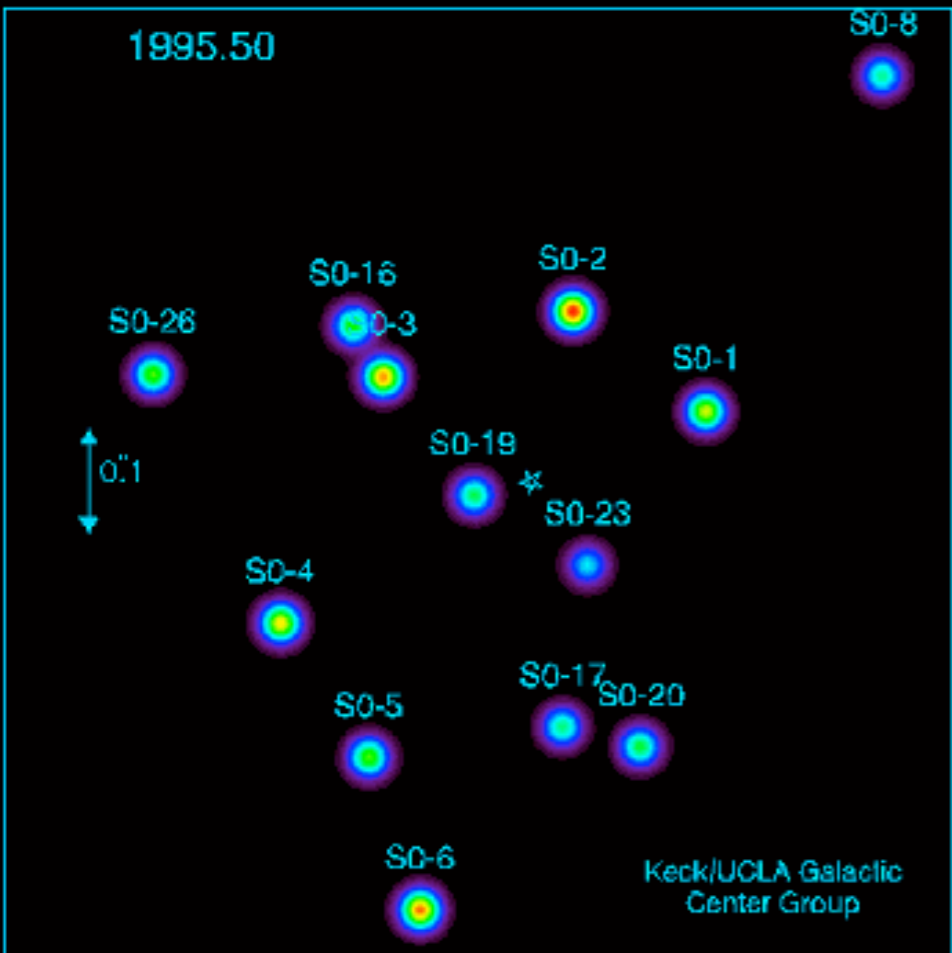




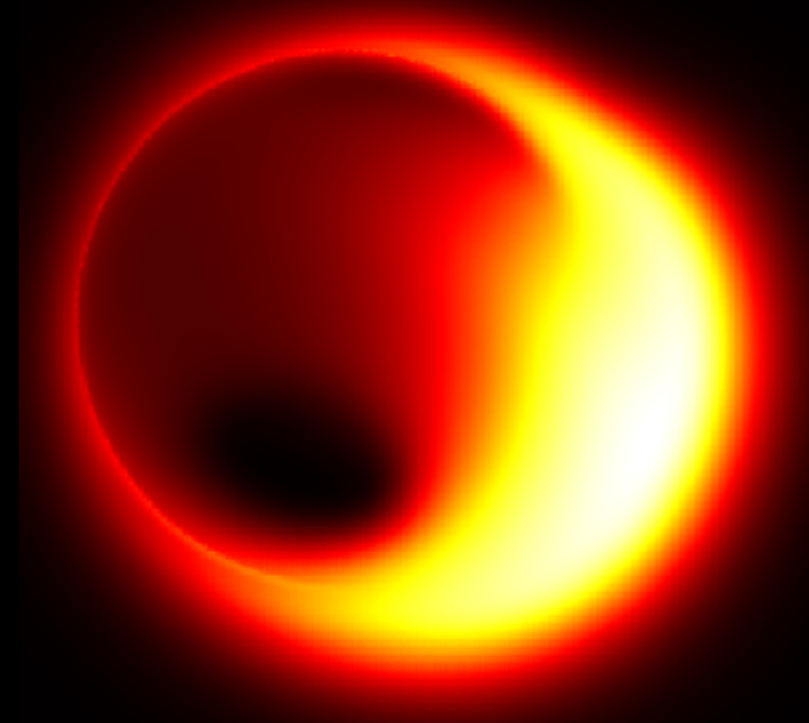
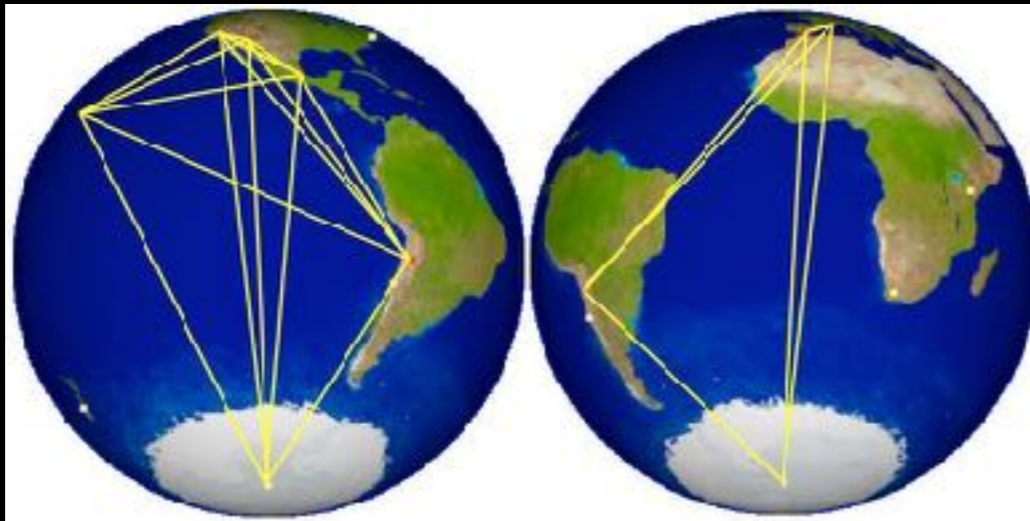
Központi fekete lyukak tömege arányos a galaxisok központi dudorának tömegével



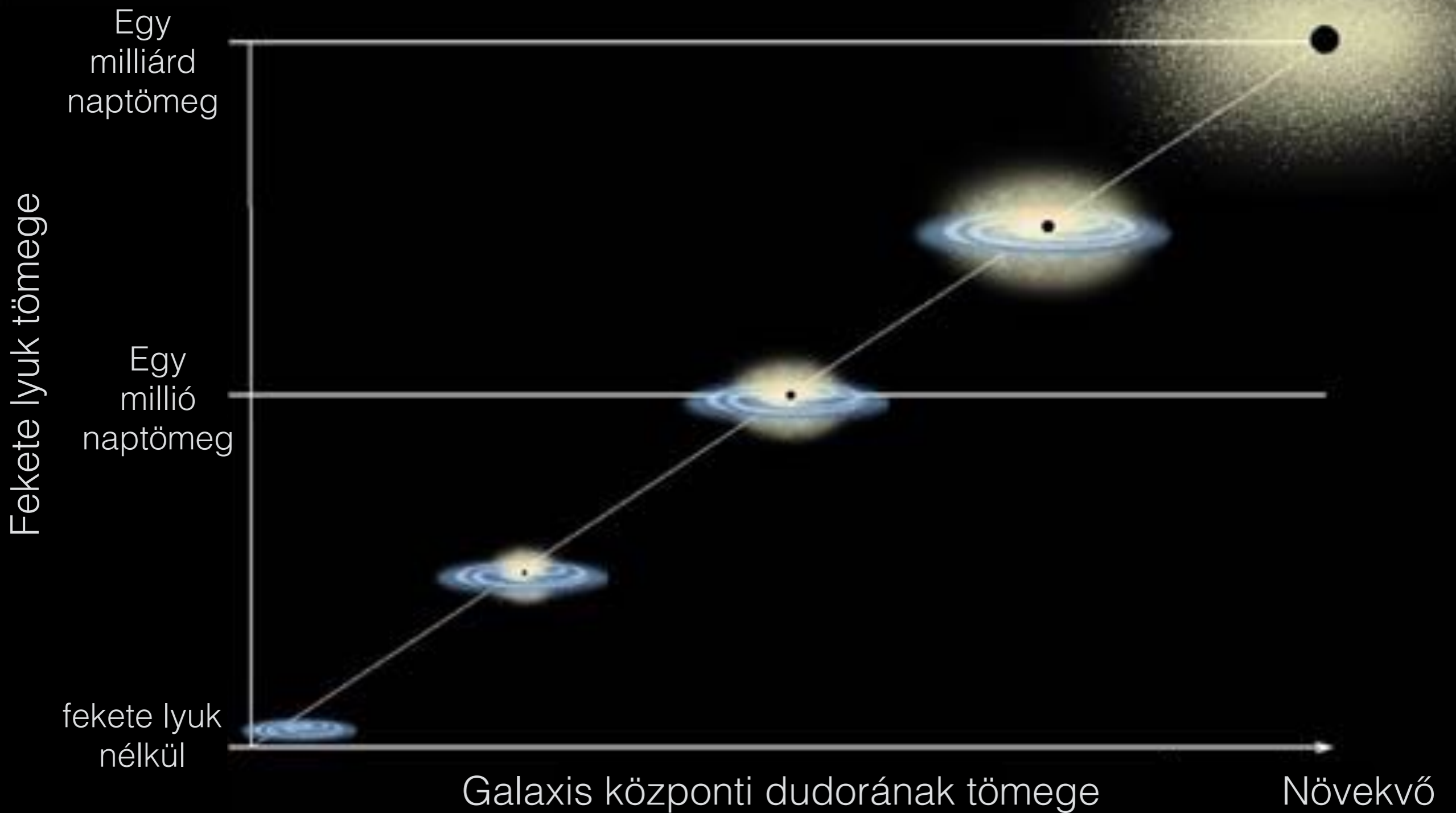




Event Horizon Telescope



Központi fekete lyukak tömege arányos a galaxisok központi dudorának tömegével



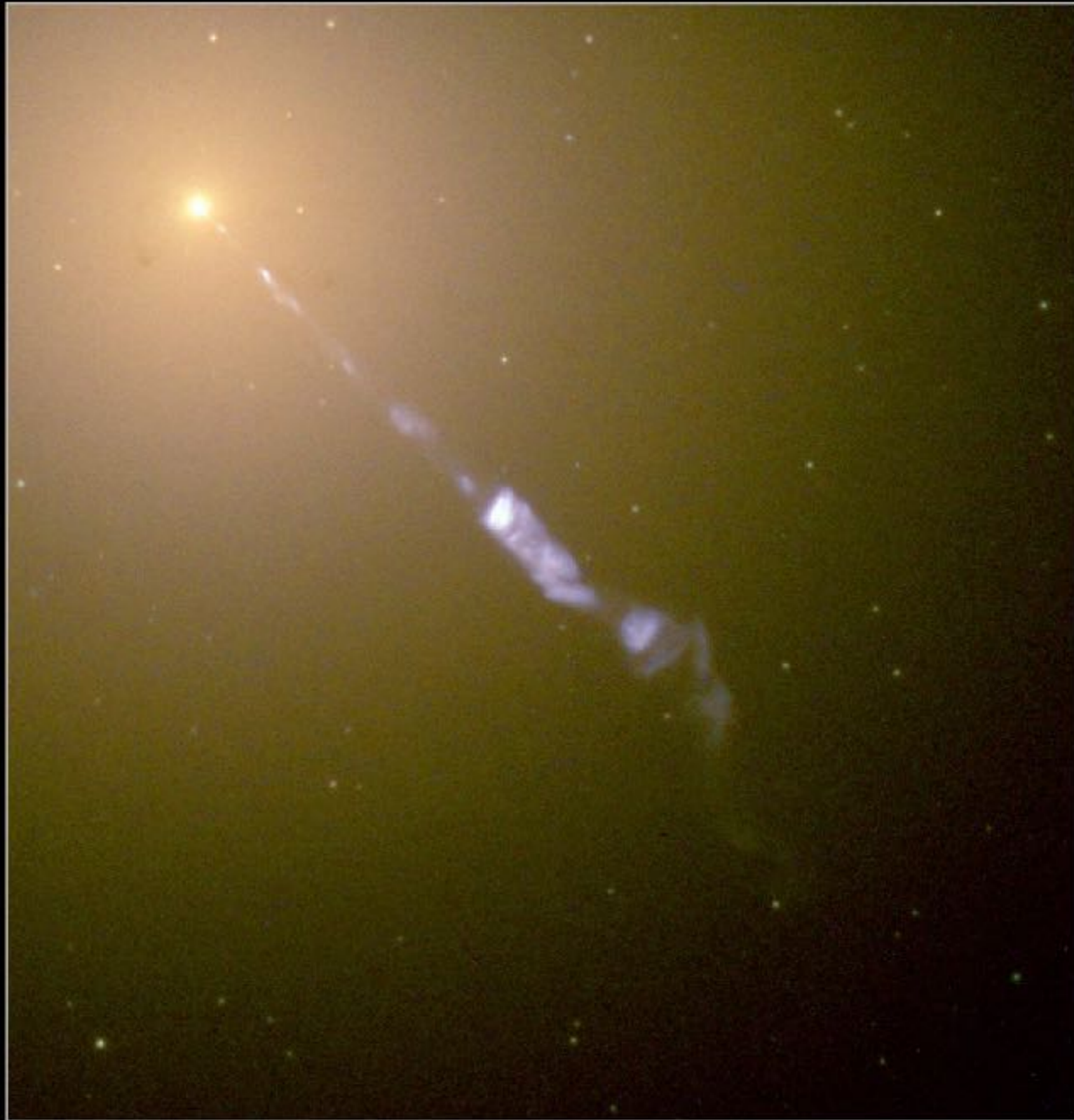


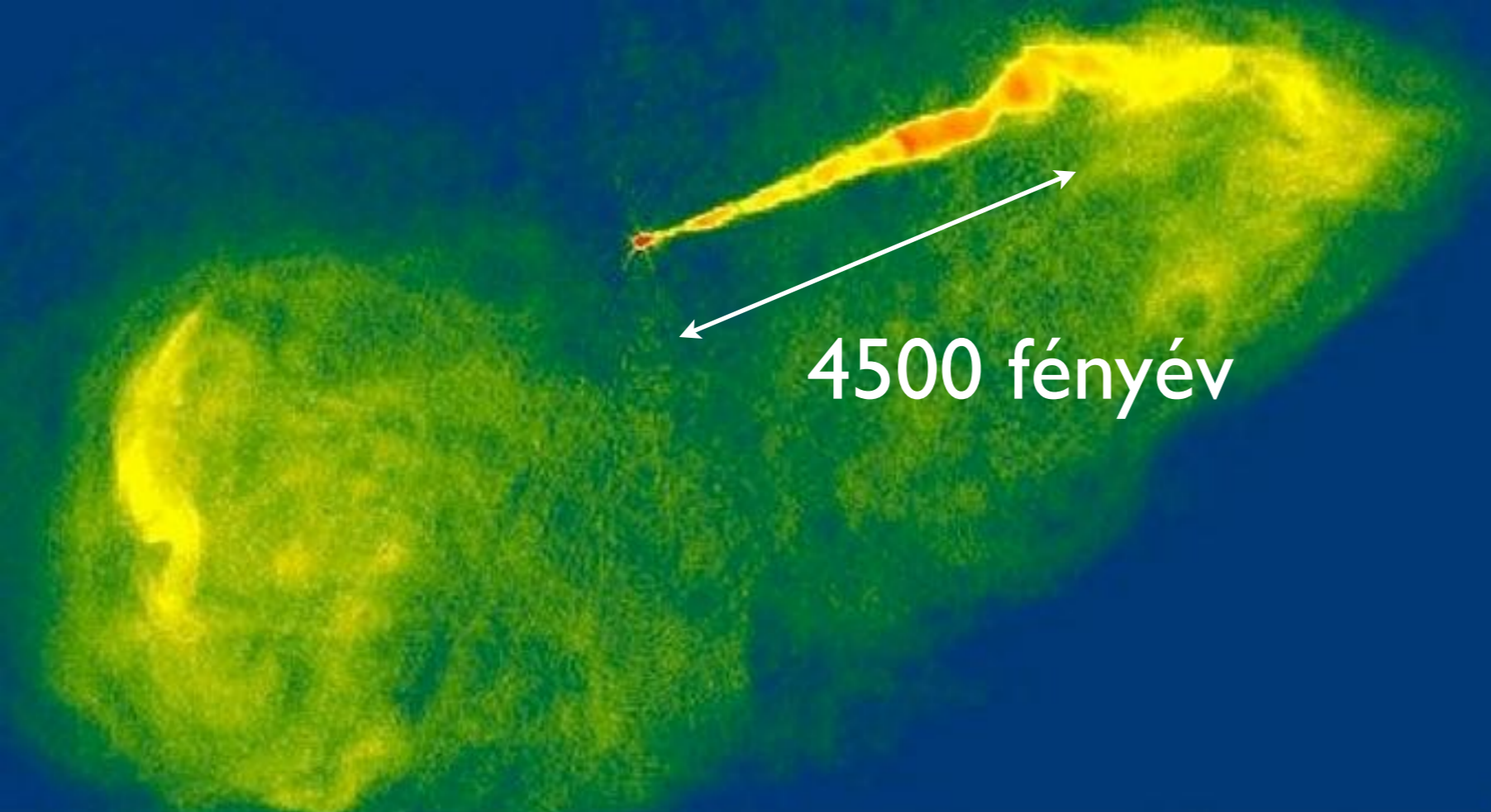




by Robert Gendler

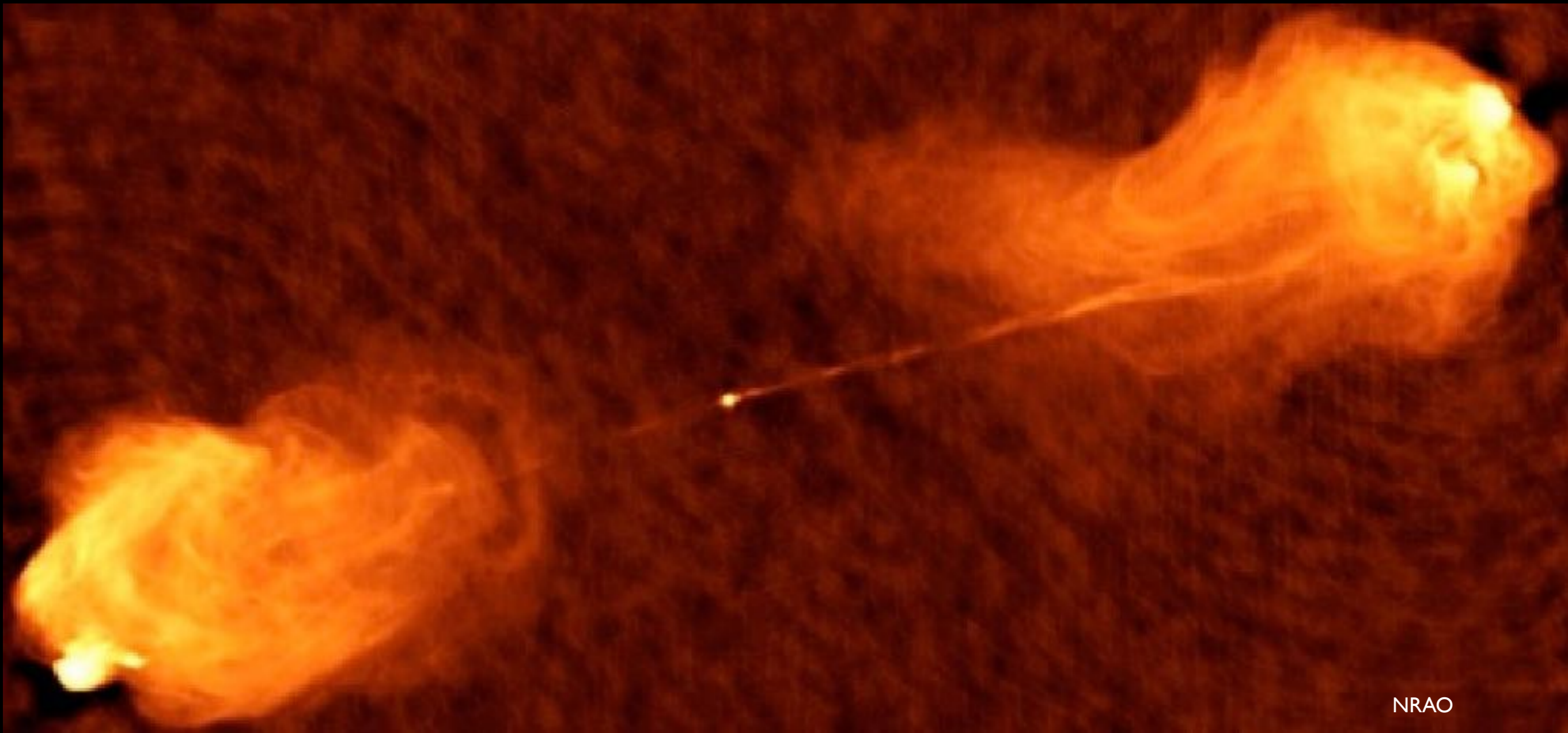
The M87 Jet





Credit:
National Radio
Astronomy
Observatory/NSF

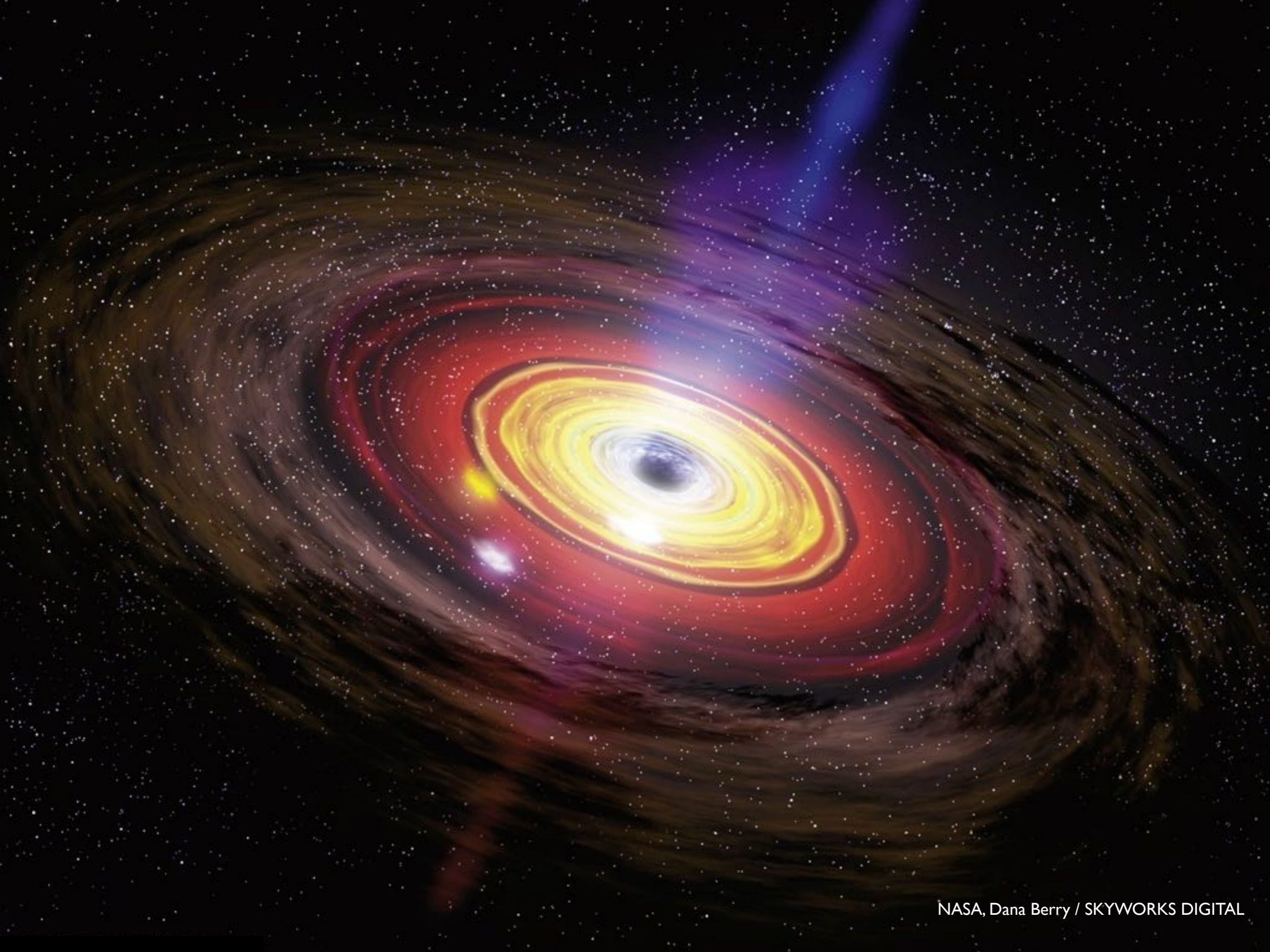
300 000 fényév!

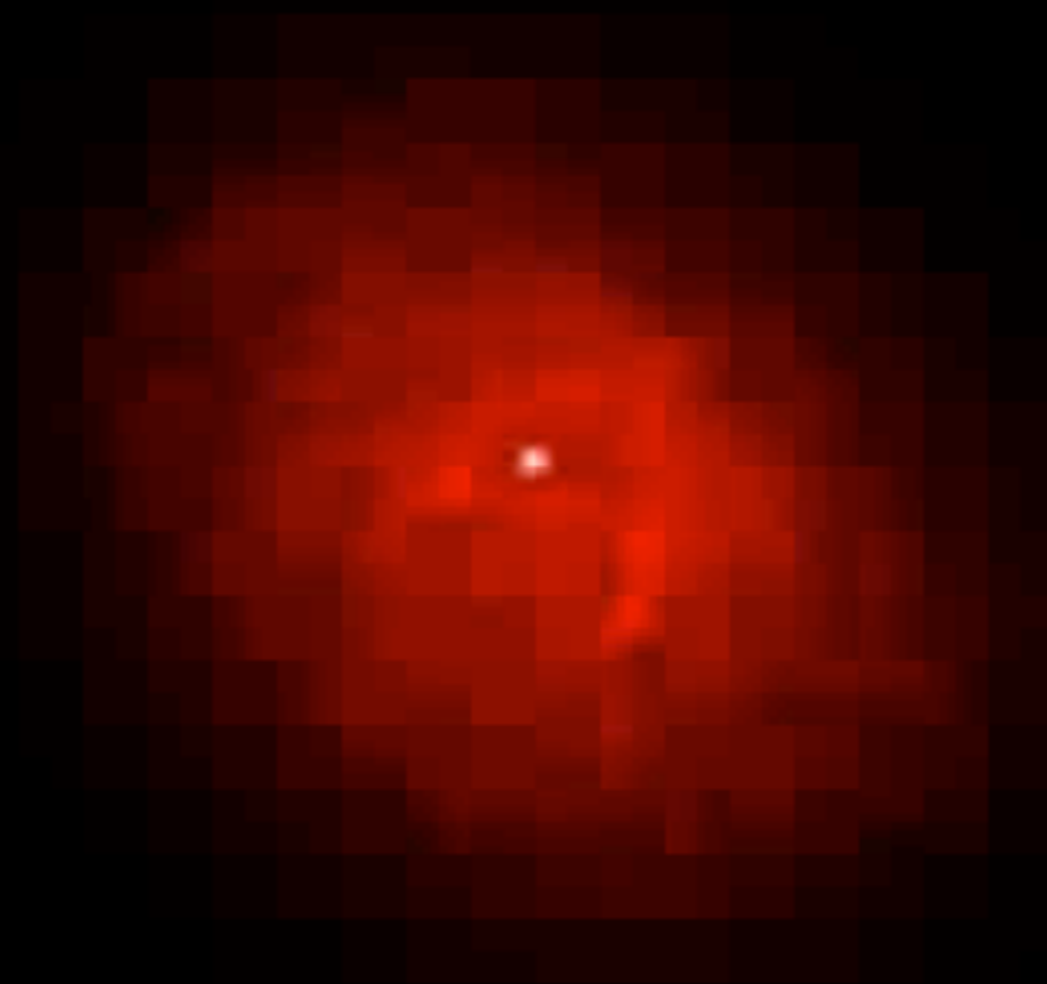


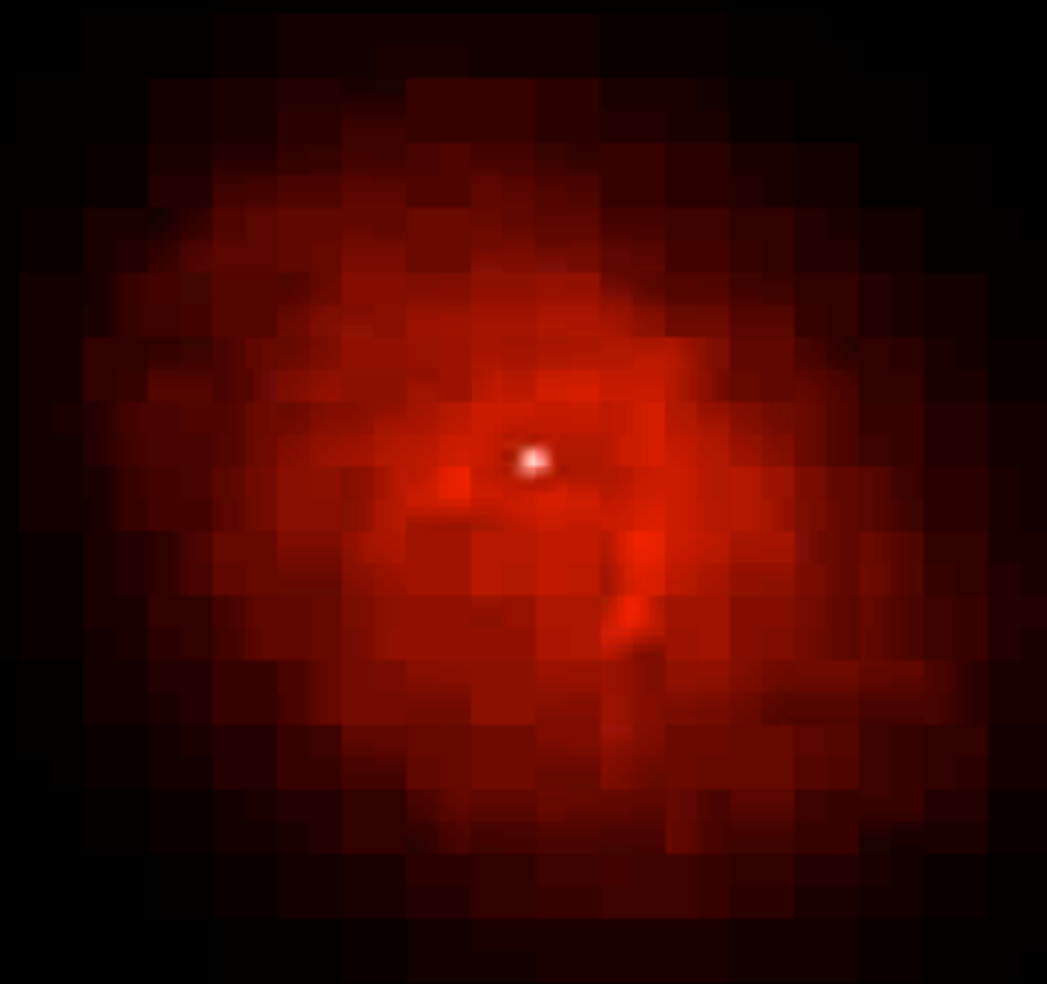
NRAO



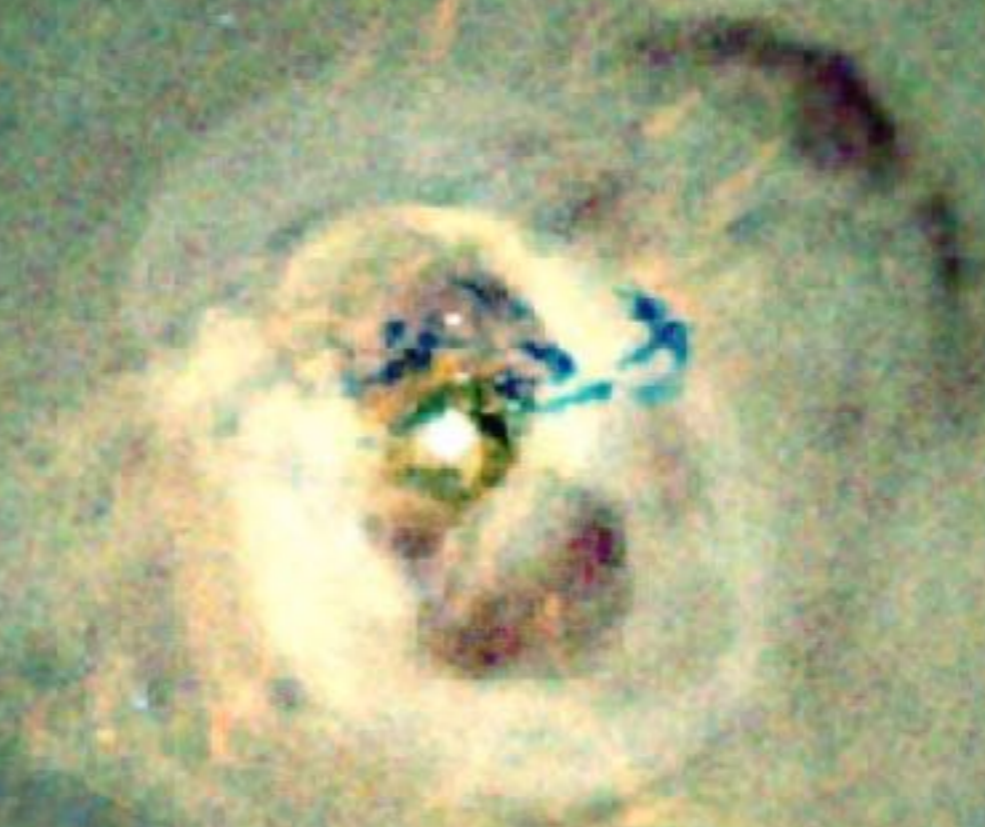
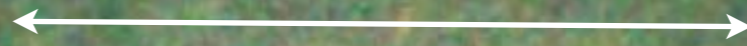








150 000 fényév



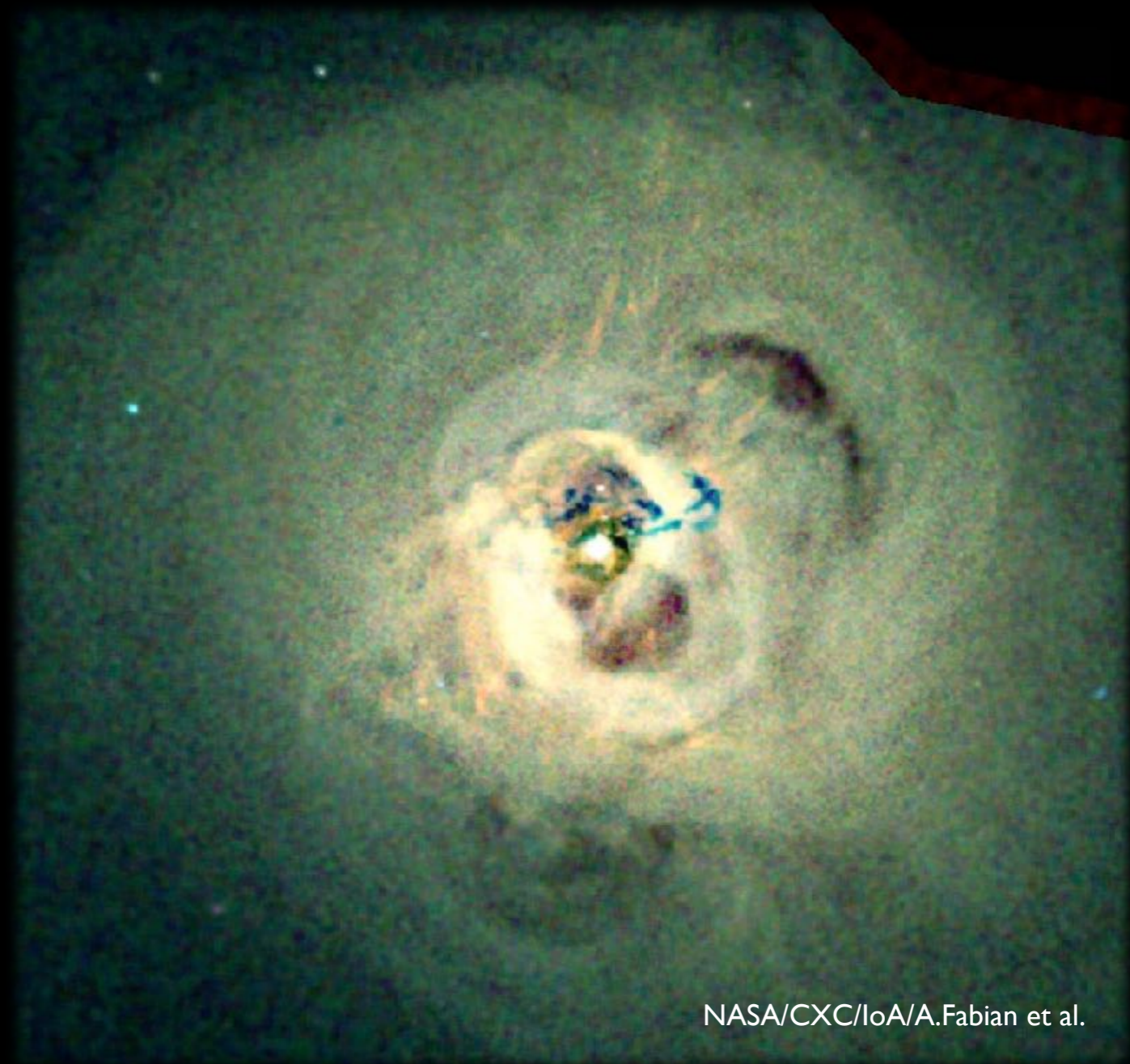
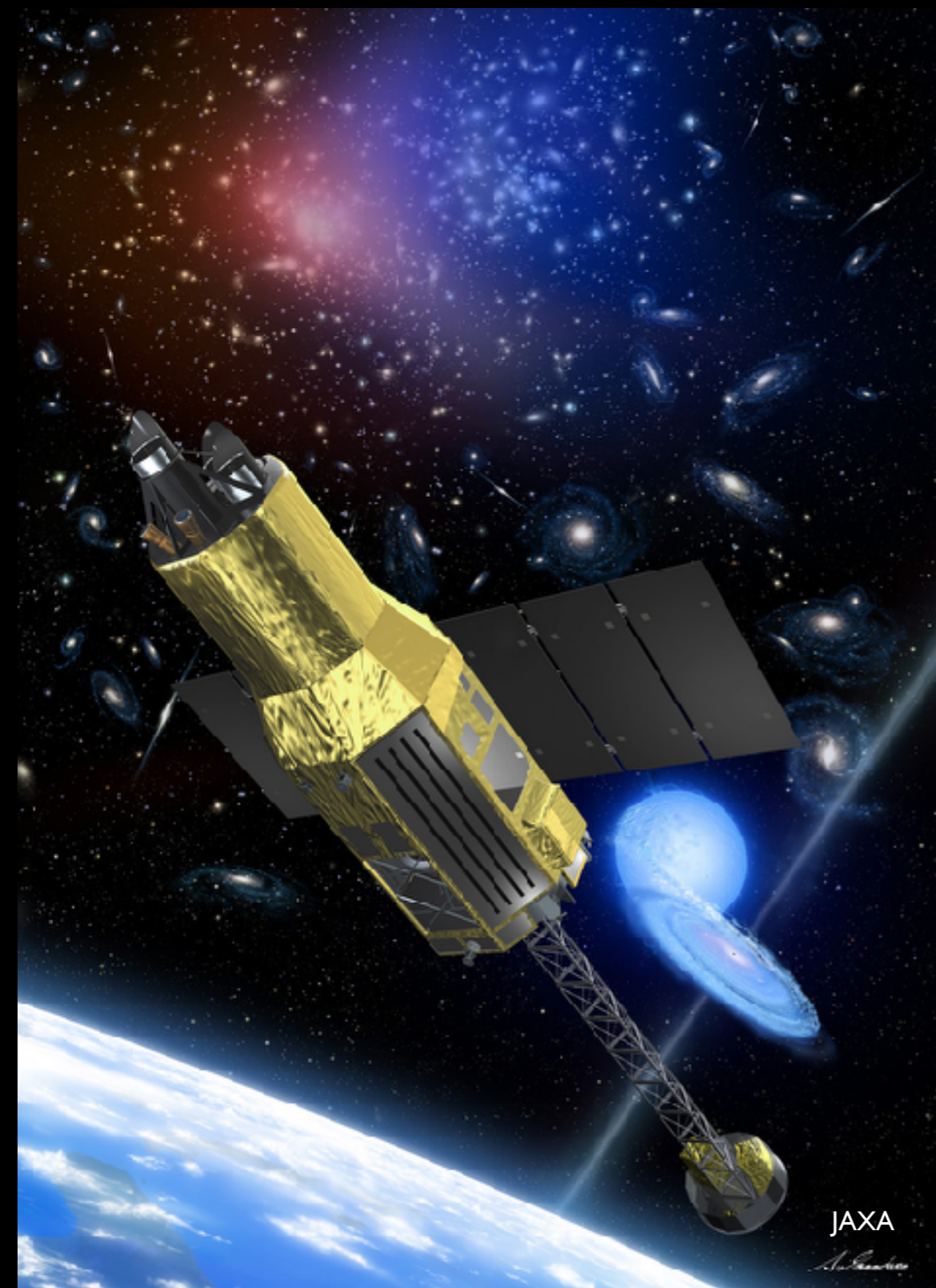
Tanegashima, Japán

Astro-H indítása, 2016 Február 17



Feb 17 - Mar 26, 2016

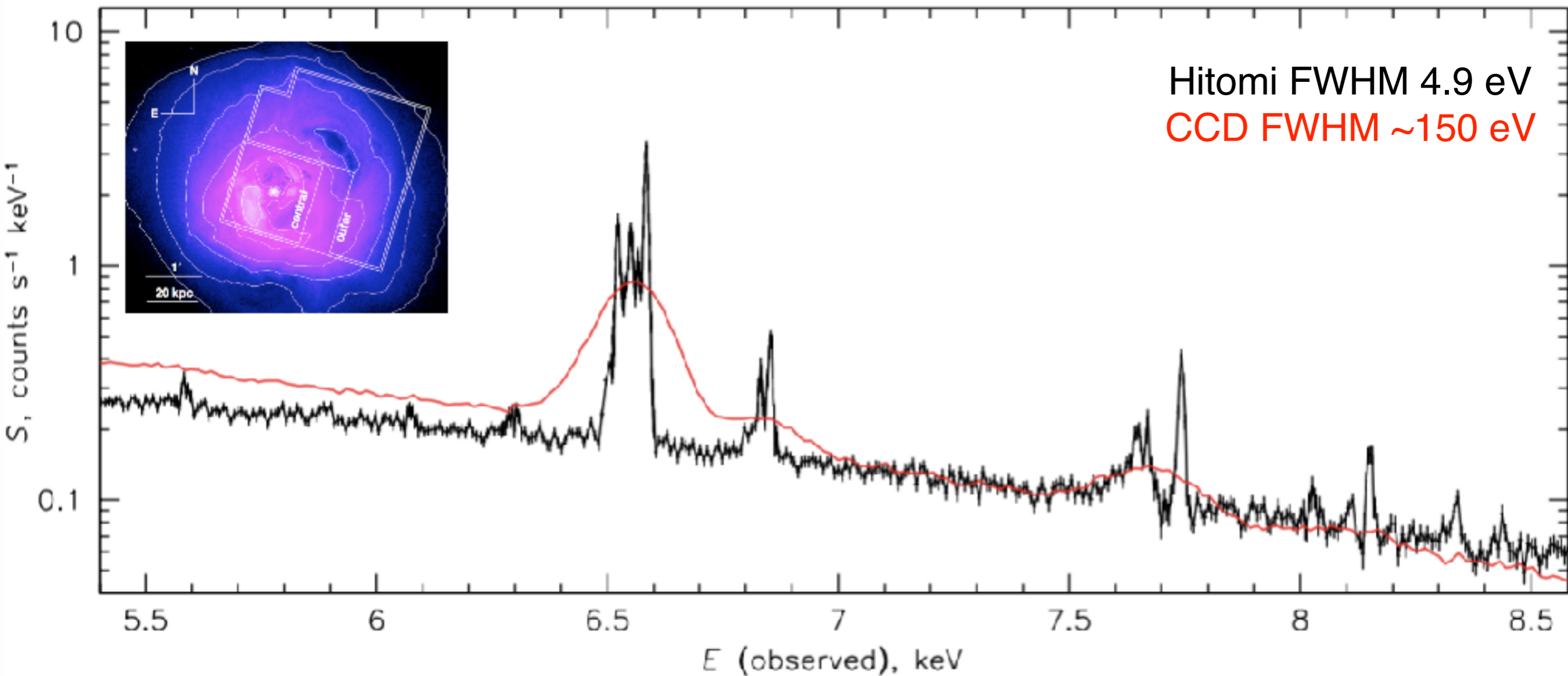
Perseus galaxishalmaz:
3 napos megfigyelés

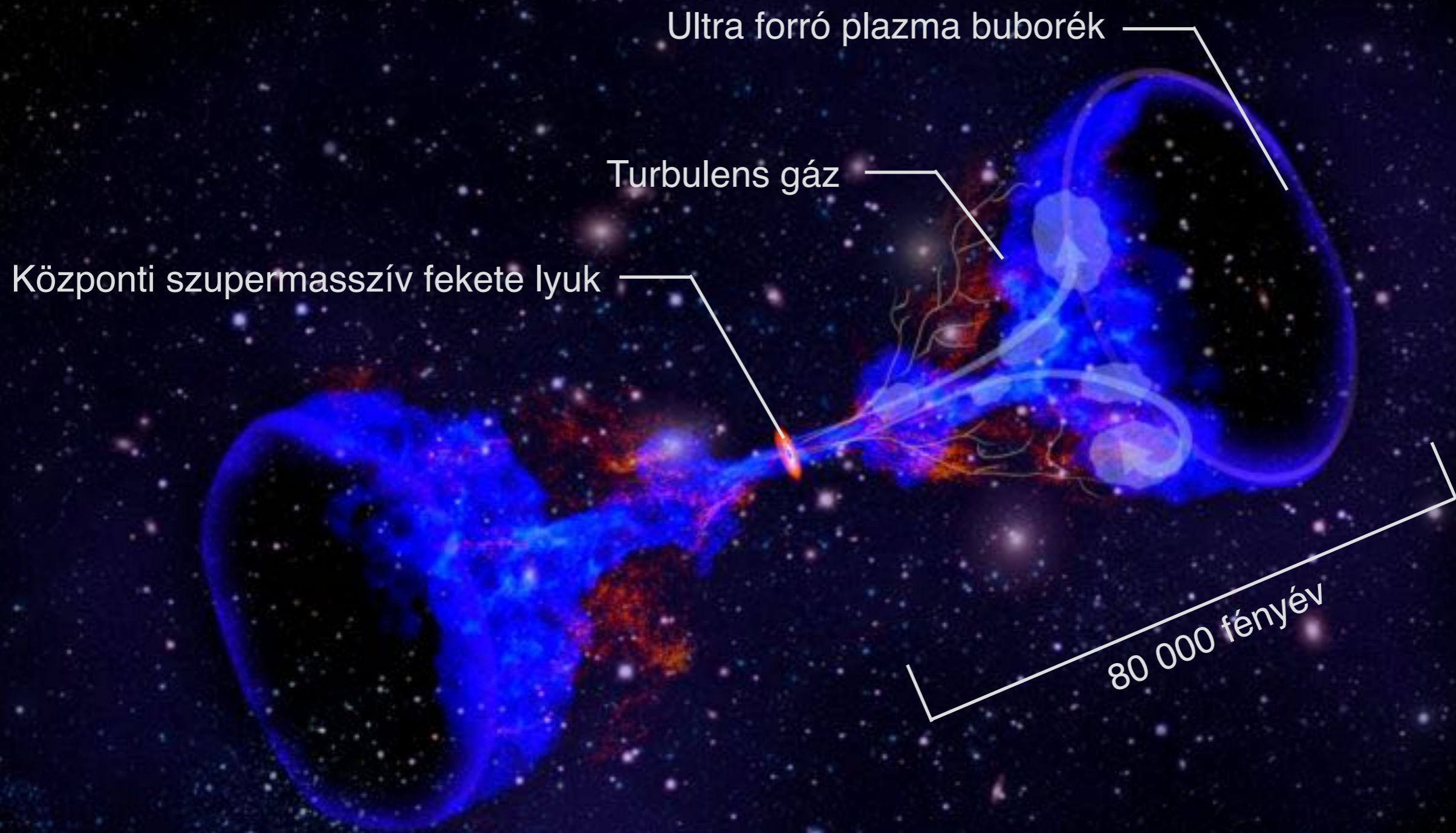


NASA/CXC/loA/A.Fabian et al.

Az első *Hitomi* (*ASTRO-H*) megfigyelés

A Perseus galaxishalmaz röntgen spektruma



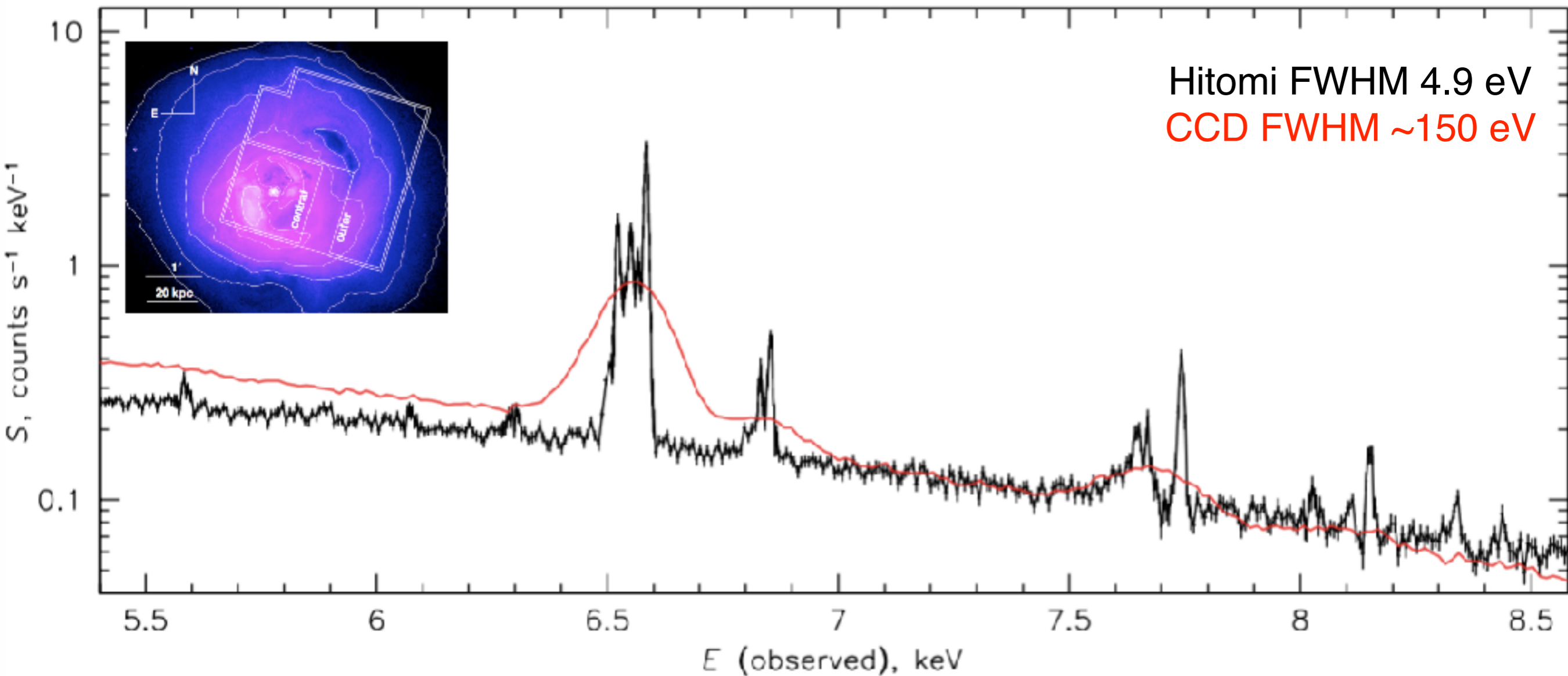






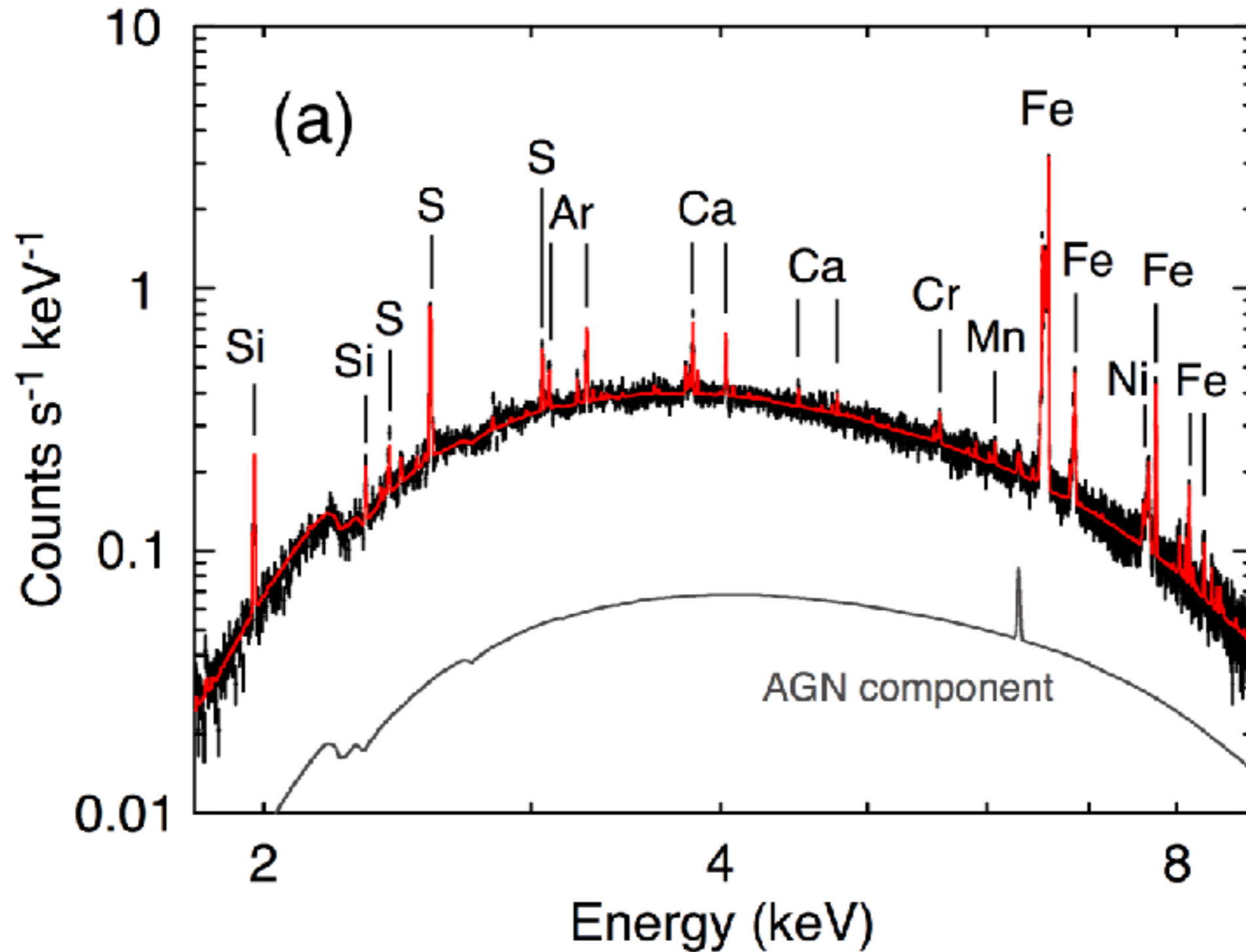
Az első *Hitomi* (*ASTRO-H*) megfigyelés

A Perseus galaxishalmaz röntgen spektruma



Az első *Hitomi* (*ASTRO-H*) megfigyelés

A Perseus galaxishalmaz röntgen spektruma



[On behalf of the *Hitomi* collaboration, Nature, submitted]
[Please do not distribute, tweet, blog these unpublished results]

A vegyi elemek eredete

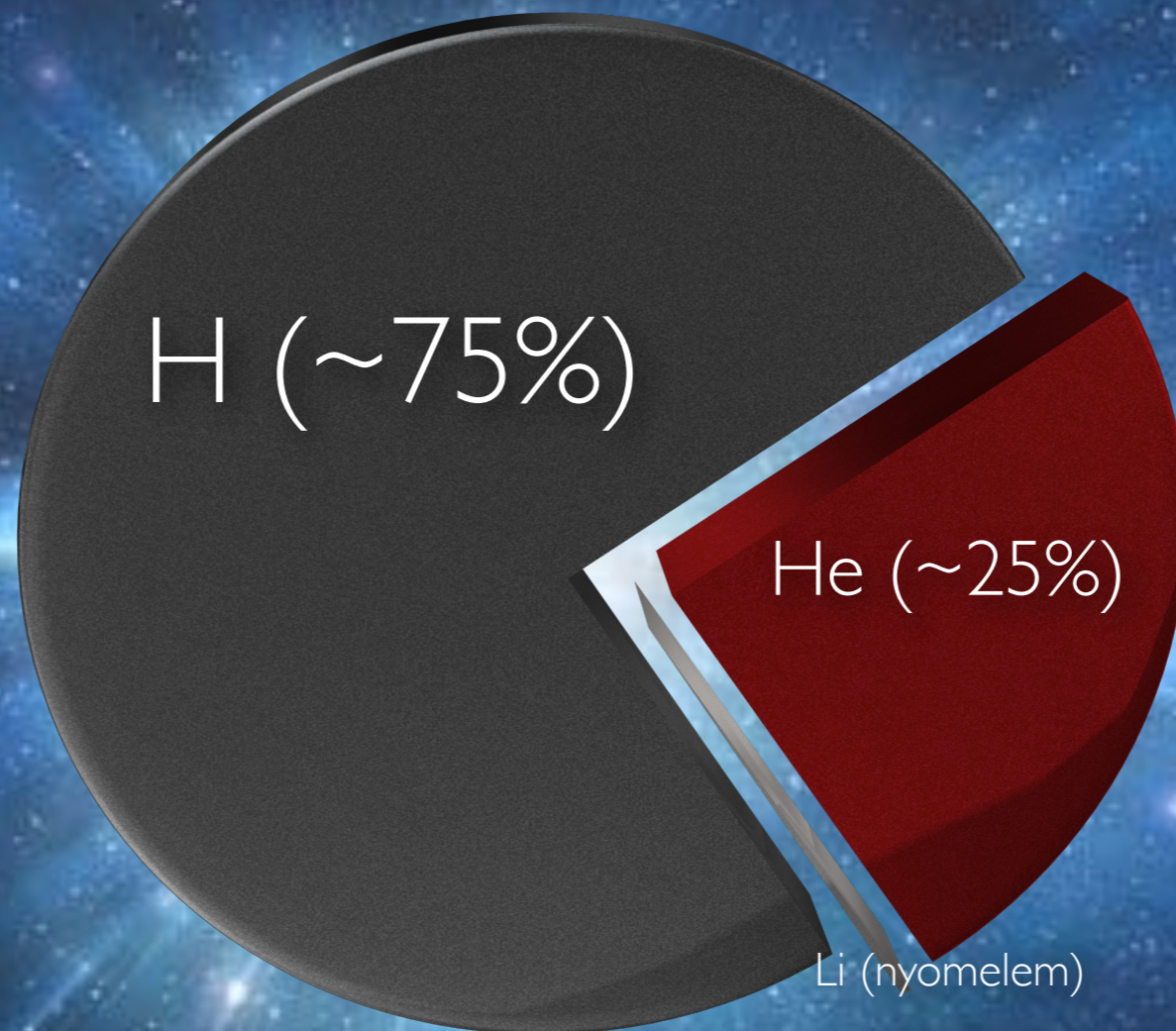
Honnan származnak a vegyi
elemek?



A vegyi elemek eredete

Honnan származnak a vegyi
elemek?

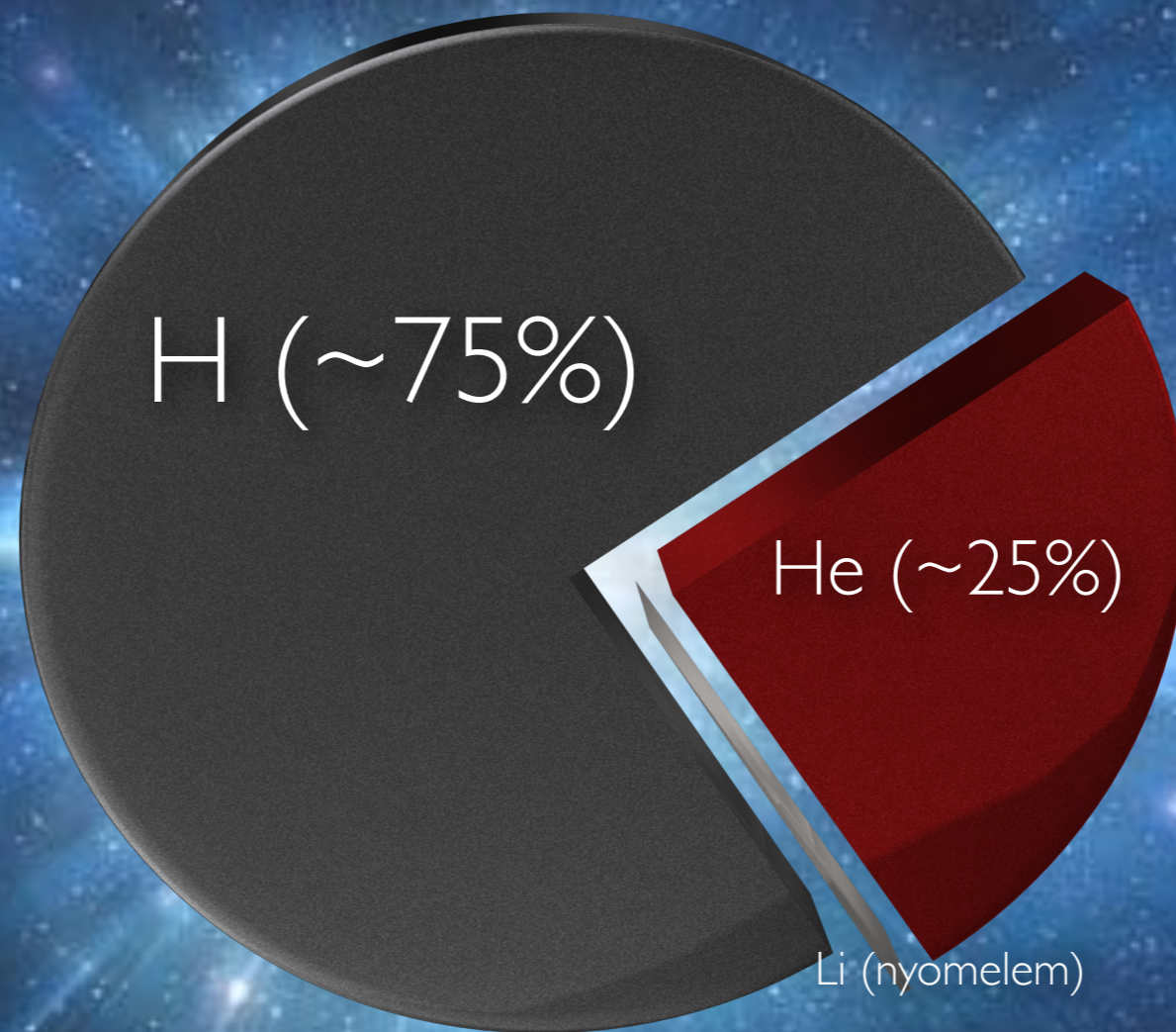
Primordiális
nukleoszintézis



A vegyi elemek eredete

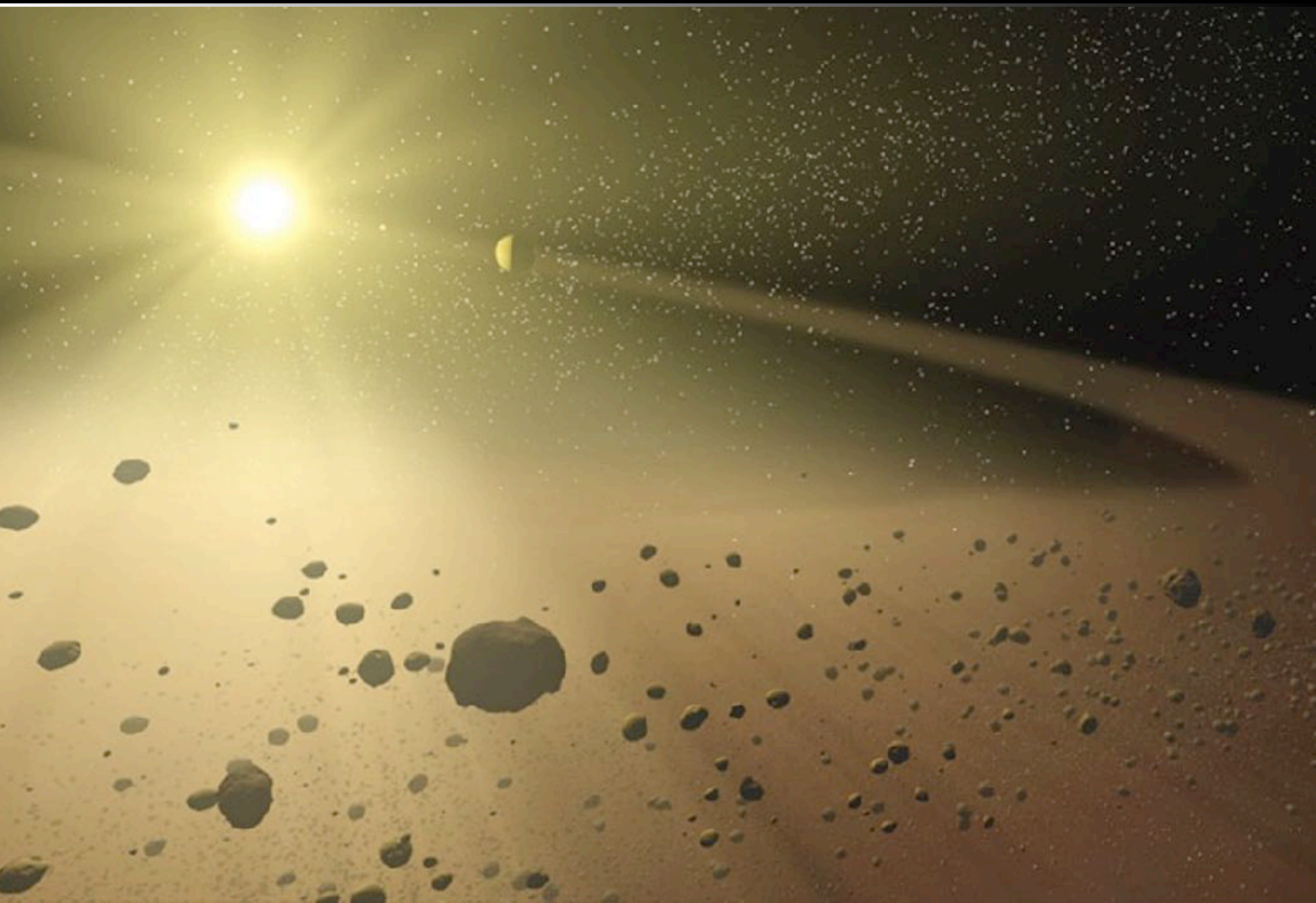
Honnan származnak a vegyi elemek?

Primordiális
nukleoszintézis

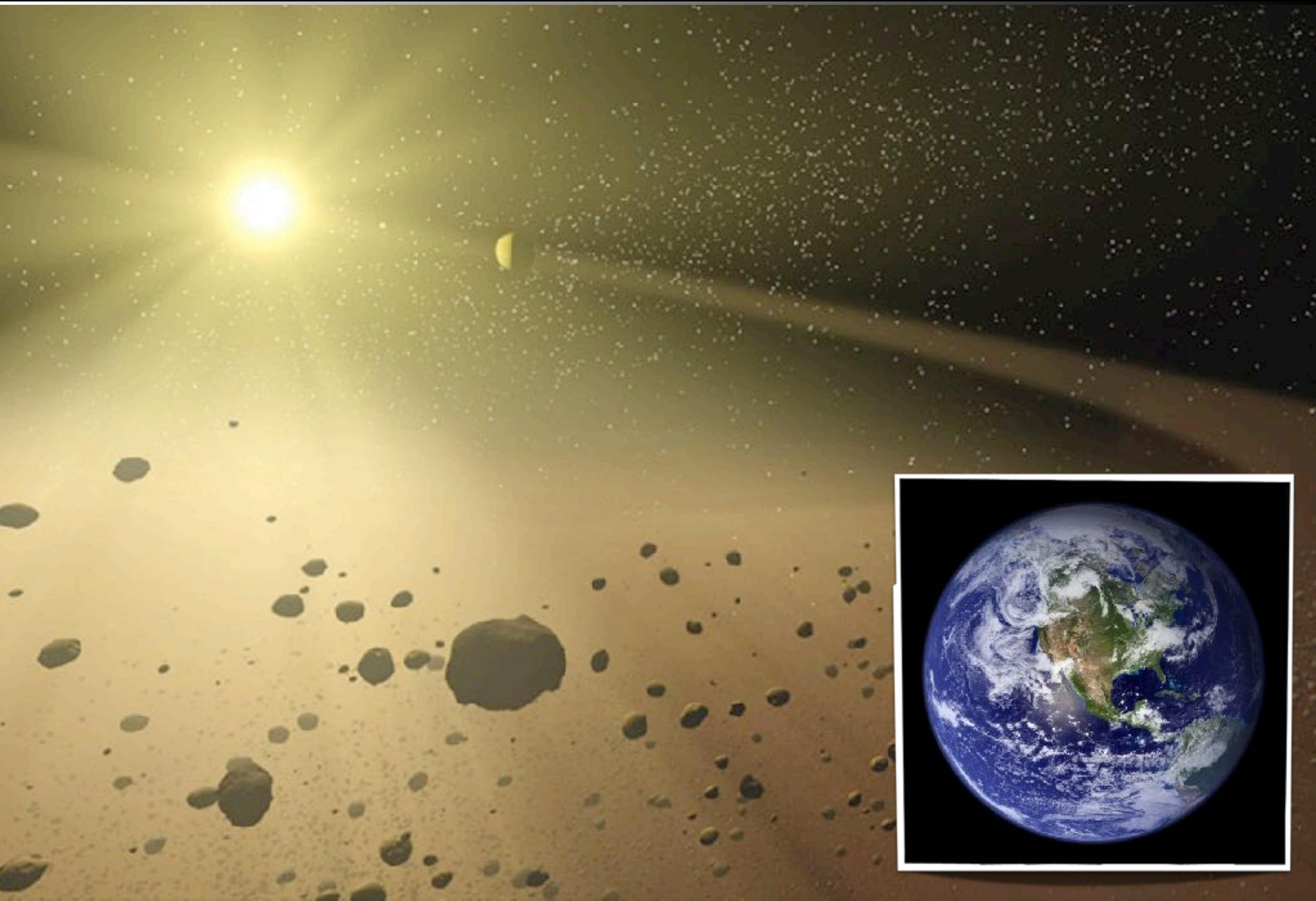


Hogyan keletkeztek a
nehezebb elemek (fémek)?

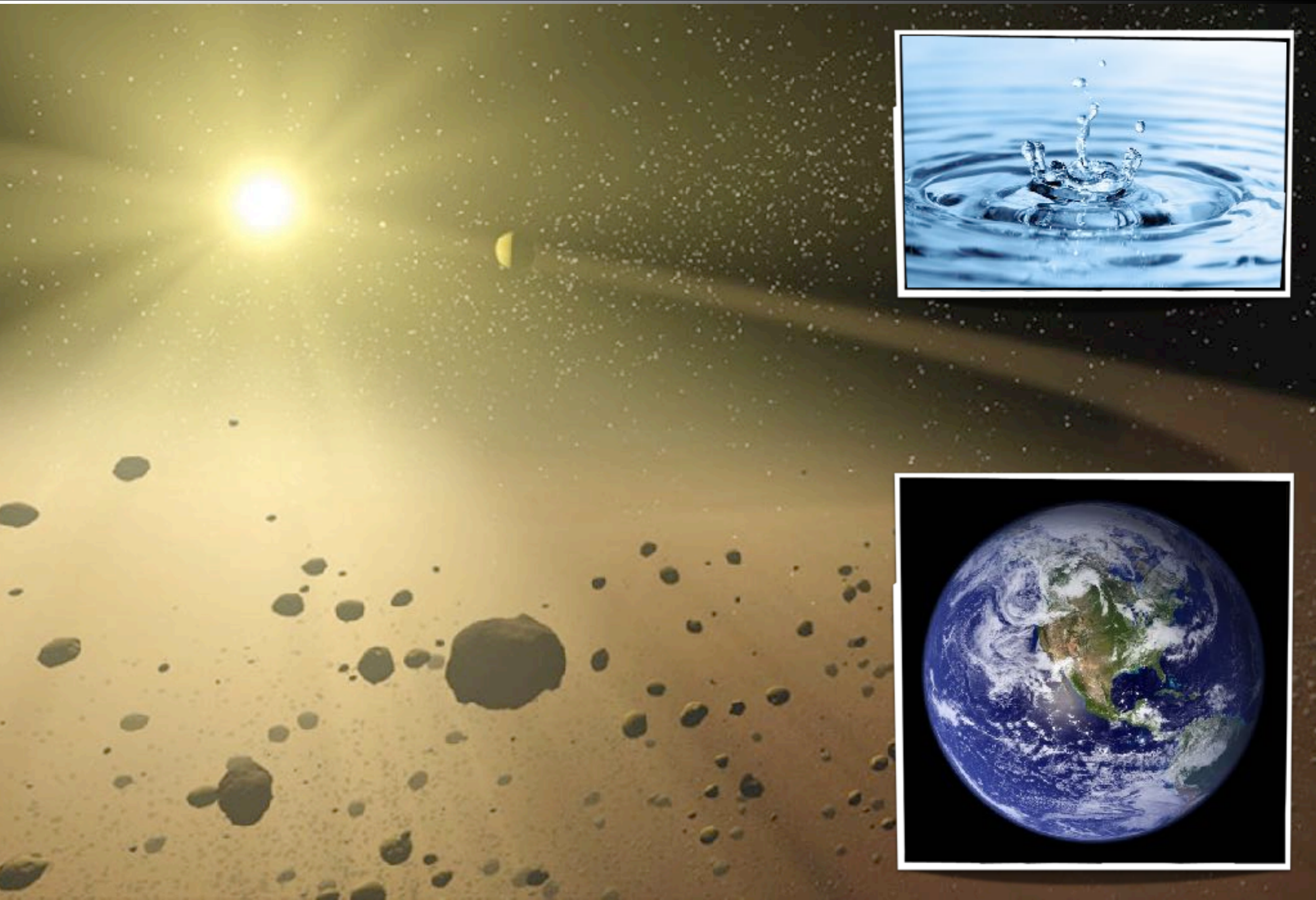
A vegyi elemek eredete



A vegyi elemek eredete



A vegyi elemek eredete



A vegyi elemek eredete



A vegyi elemek eredete



A vegyi elemek eredete



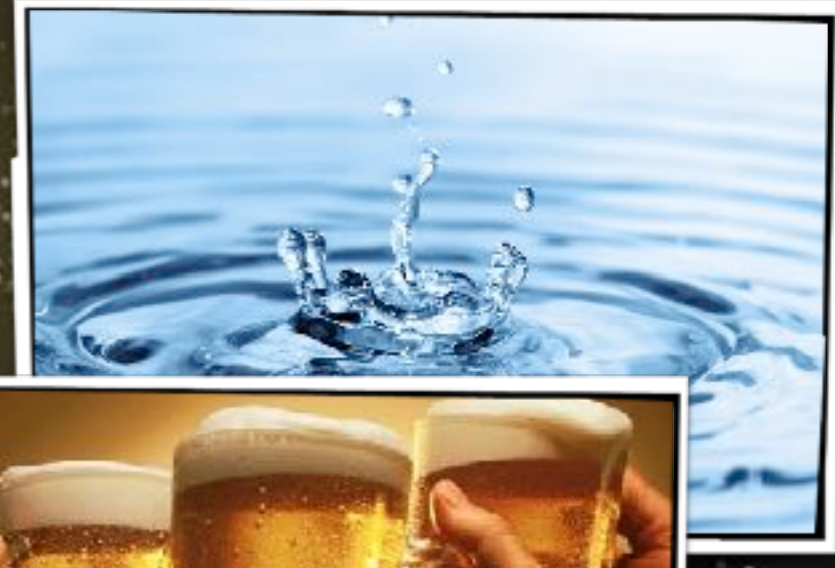
A vegyi elemek eredete



A vegyi elemek eredete



A vegyi elemek eredete



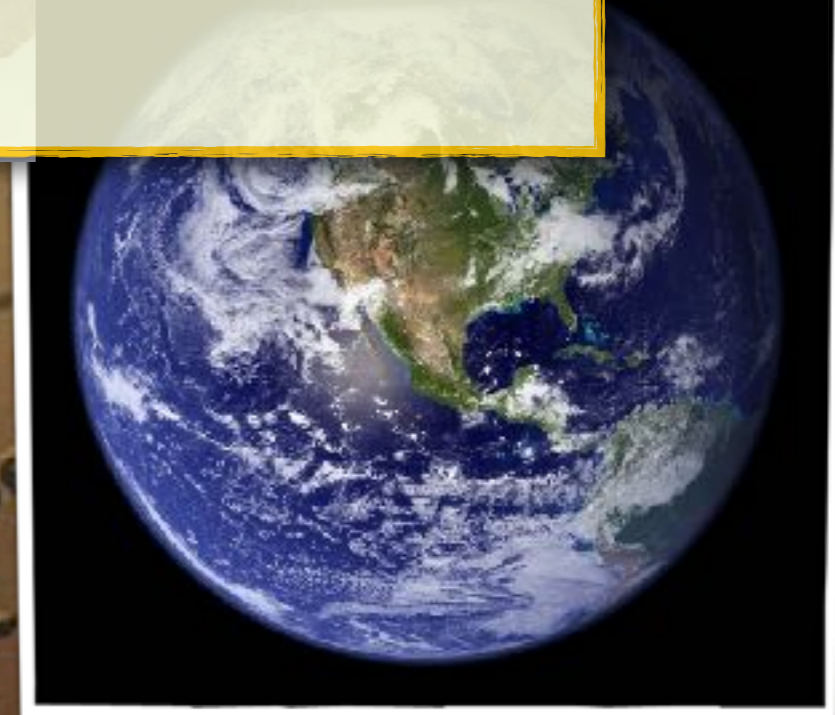
A vegyi elemek eredete



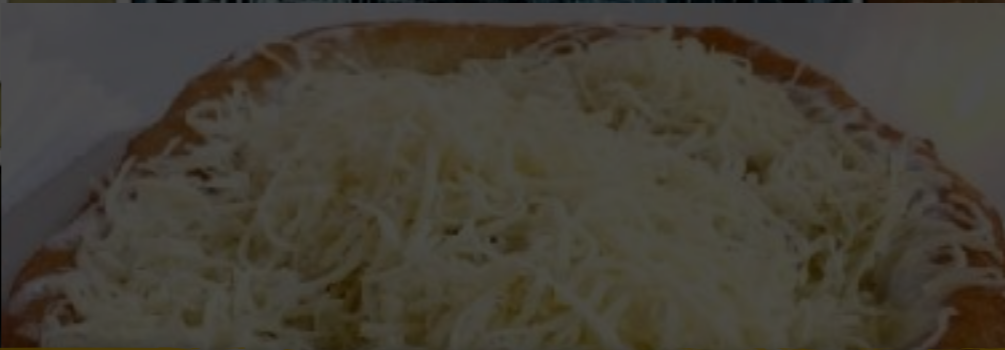
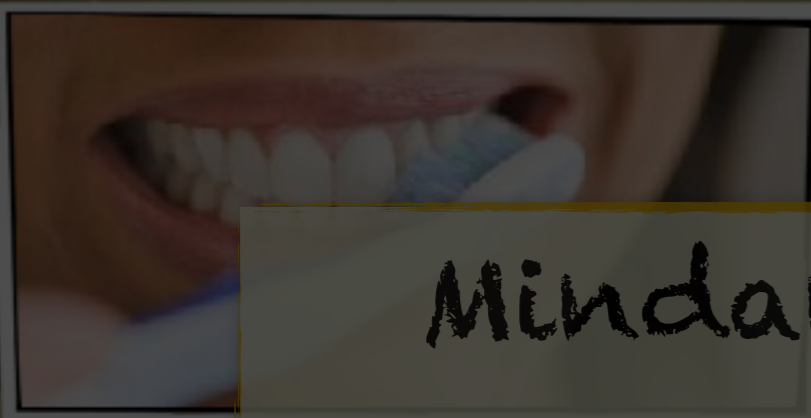
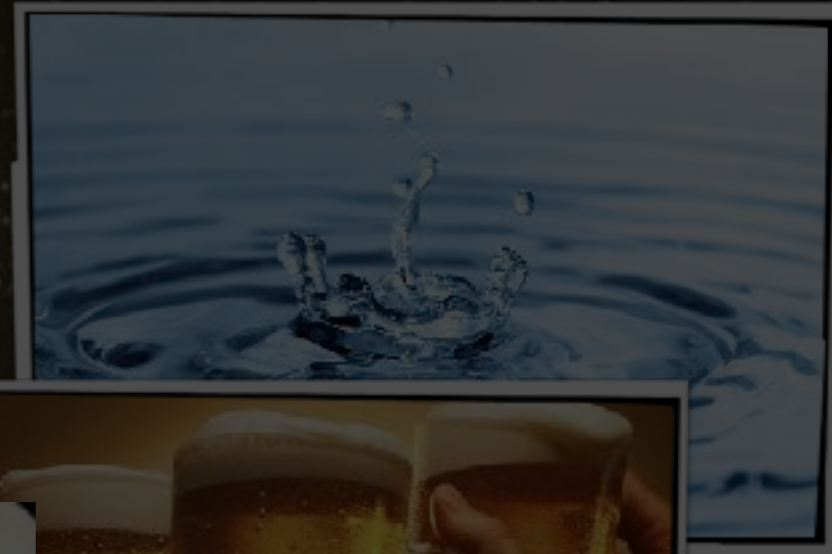
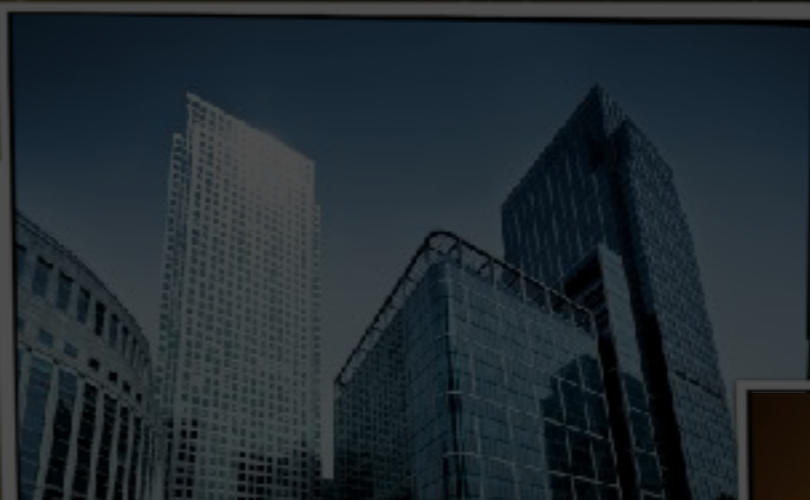
A vegyi elemek eredete



Mindannyian a csillagokból
származunk!

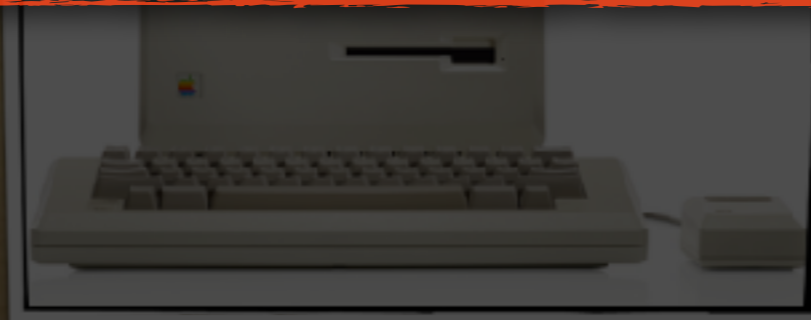
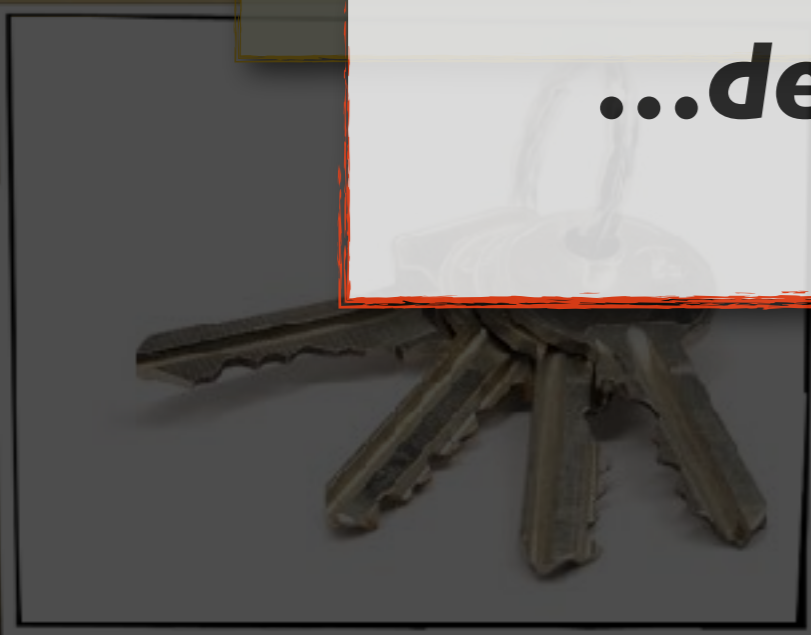


A vegyi elemek eredete

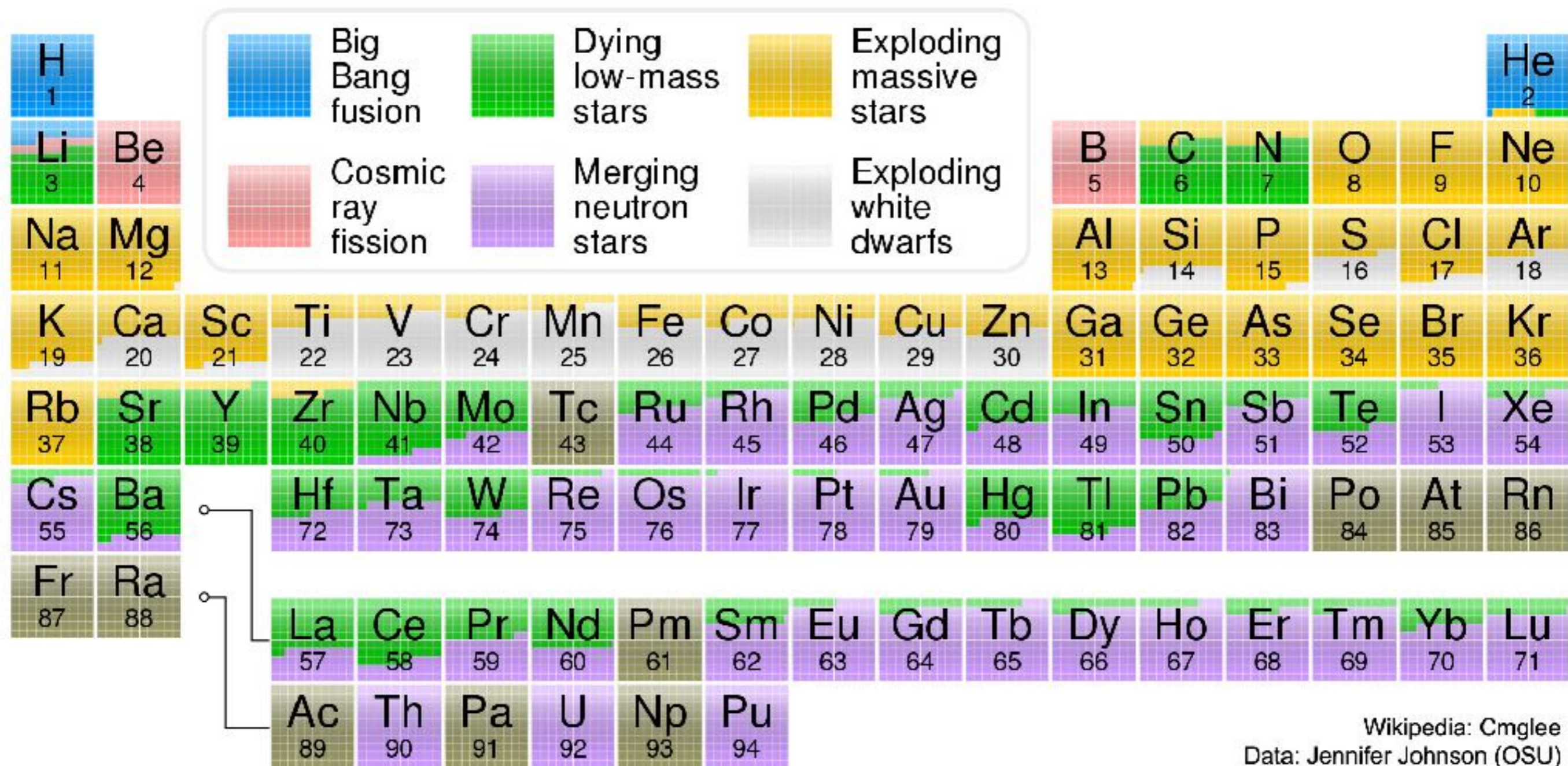


Mindannyian a csillagokból
származunk!

**...de a szupermóvákat nem
értjük...**



A vegyi elemek eredete



Wikipedia: Cmglee
Data: Jennifer Johnson (OSU)

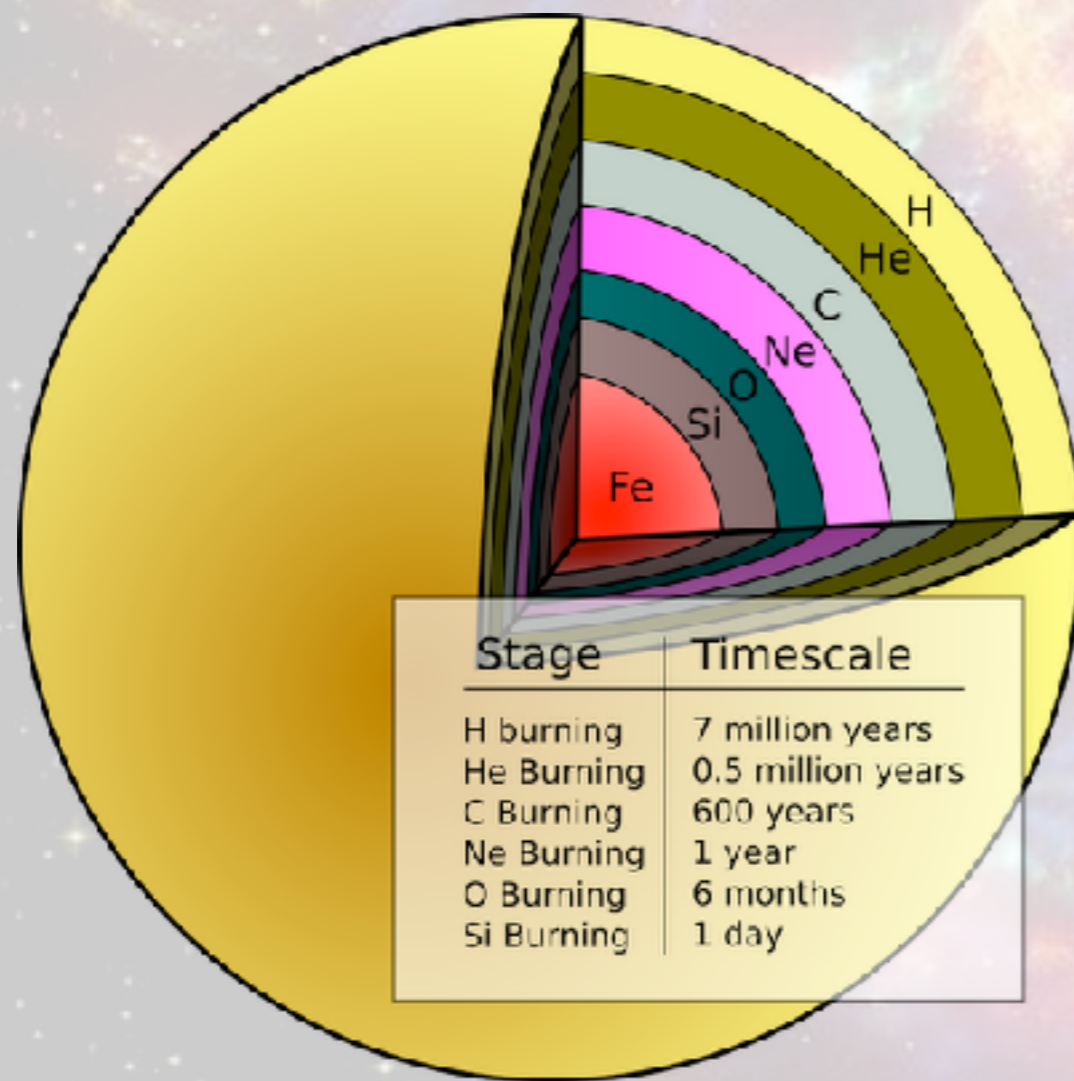
A vegyi elemek eredete

1) Mag összeomlás szupernóvák



A vegyi elemek eredete

1) Mag összeomlás szupernóvák



Termékek:

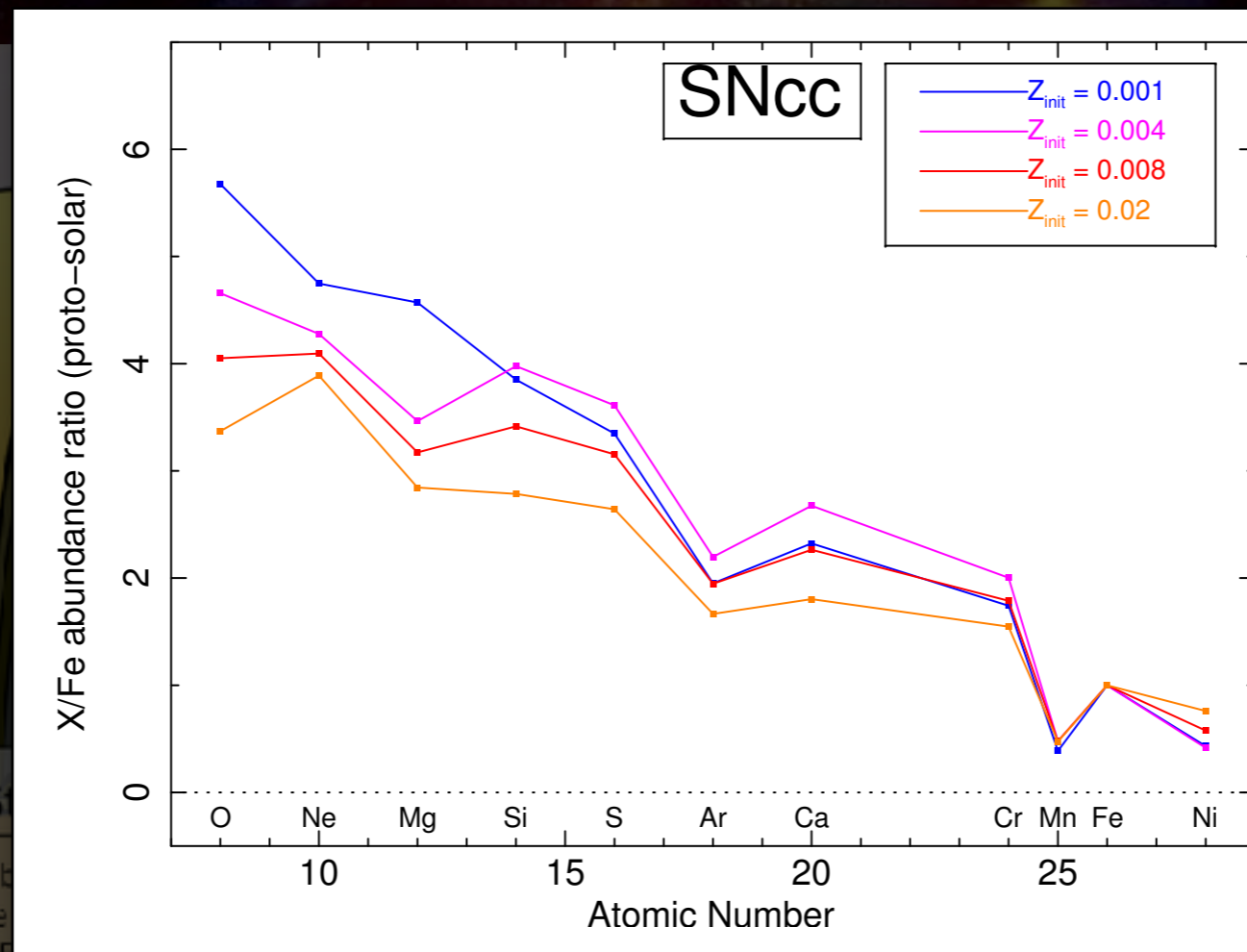
- ➔ O
- ➔ Ne
- ➔ Mg
- ➔ Si

Mennyi? Több mindentől függ:

- ➔ A csillagok tömegétől
- ➔ A csillagok fémtartalmától

A vegyi elemek eredete

1) Mag összeomlás szupernóvák



Ne Burning	1 year
O Burning	6 months
Si Burning	1 day

lásd Nomoto et al. (2013)

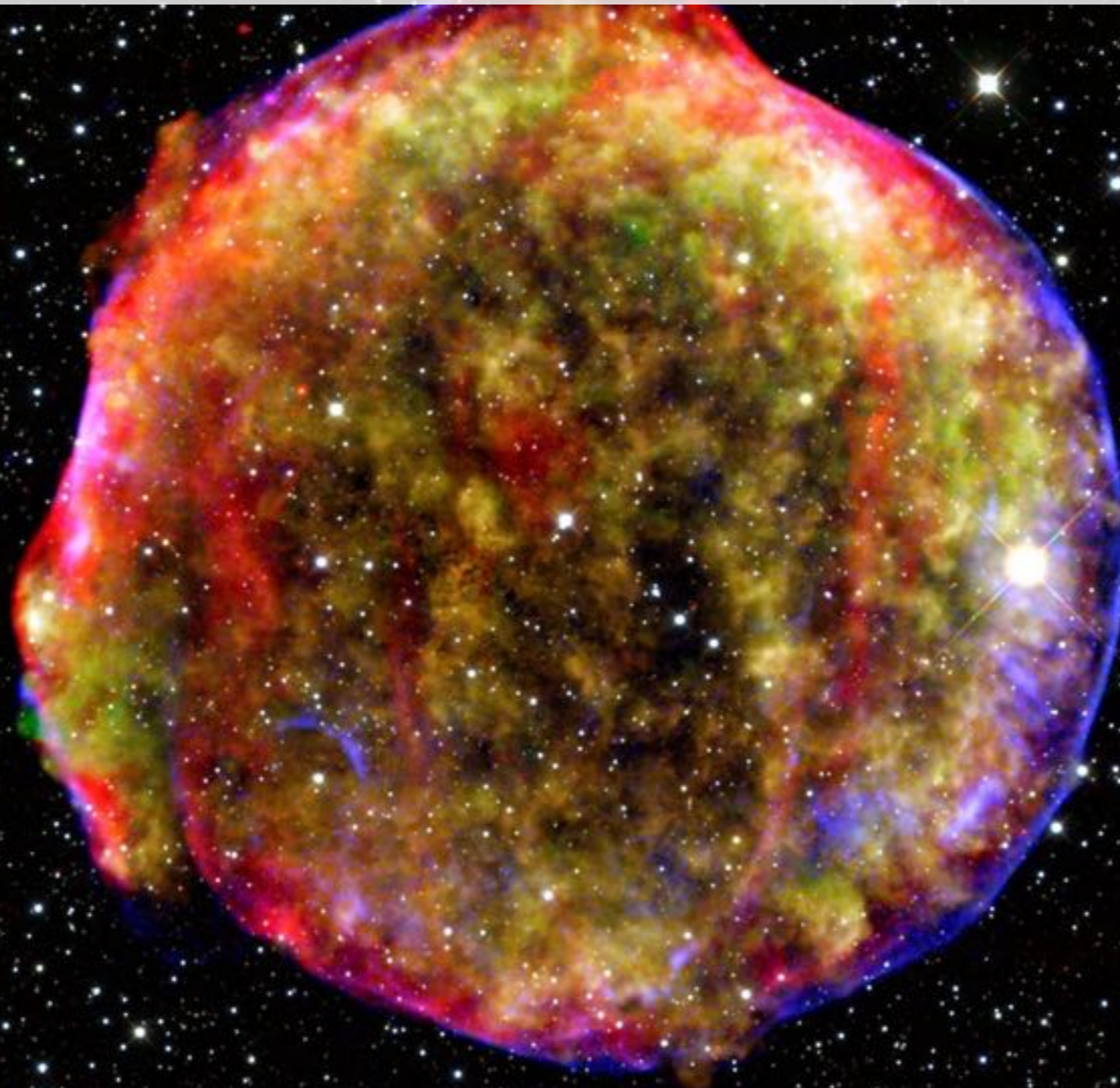
➡ *Meennyi? Iobb mindentől függ:*

➡ A csillagok tömegétől

➡ A csillagok fémtartalmától

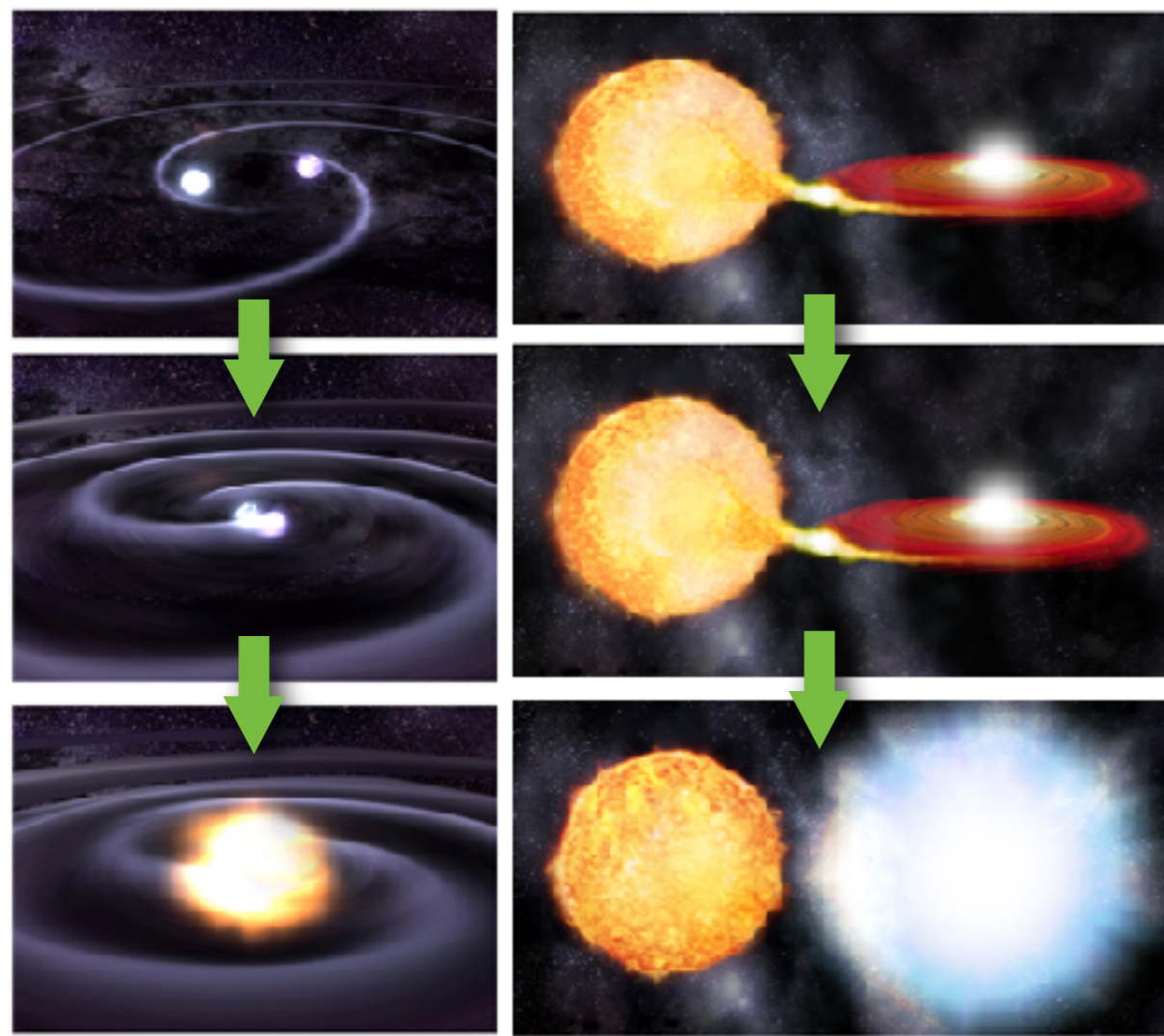
A vegyi elemek eredete

2) Ia típusú (termionukleáris) szupernóvák



A vegyi elemek eredete

2) Ia típusú (termionukleáris) szupernóvák



2 fehér
törpe?

1 fehér
törpe + csillag

Termékek:

➔ *Si*

➔ *Ca*

➔ *S*

➔ *Fe*

➔ *Ar*

➔ *Ni*

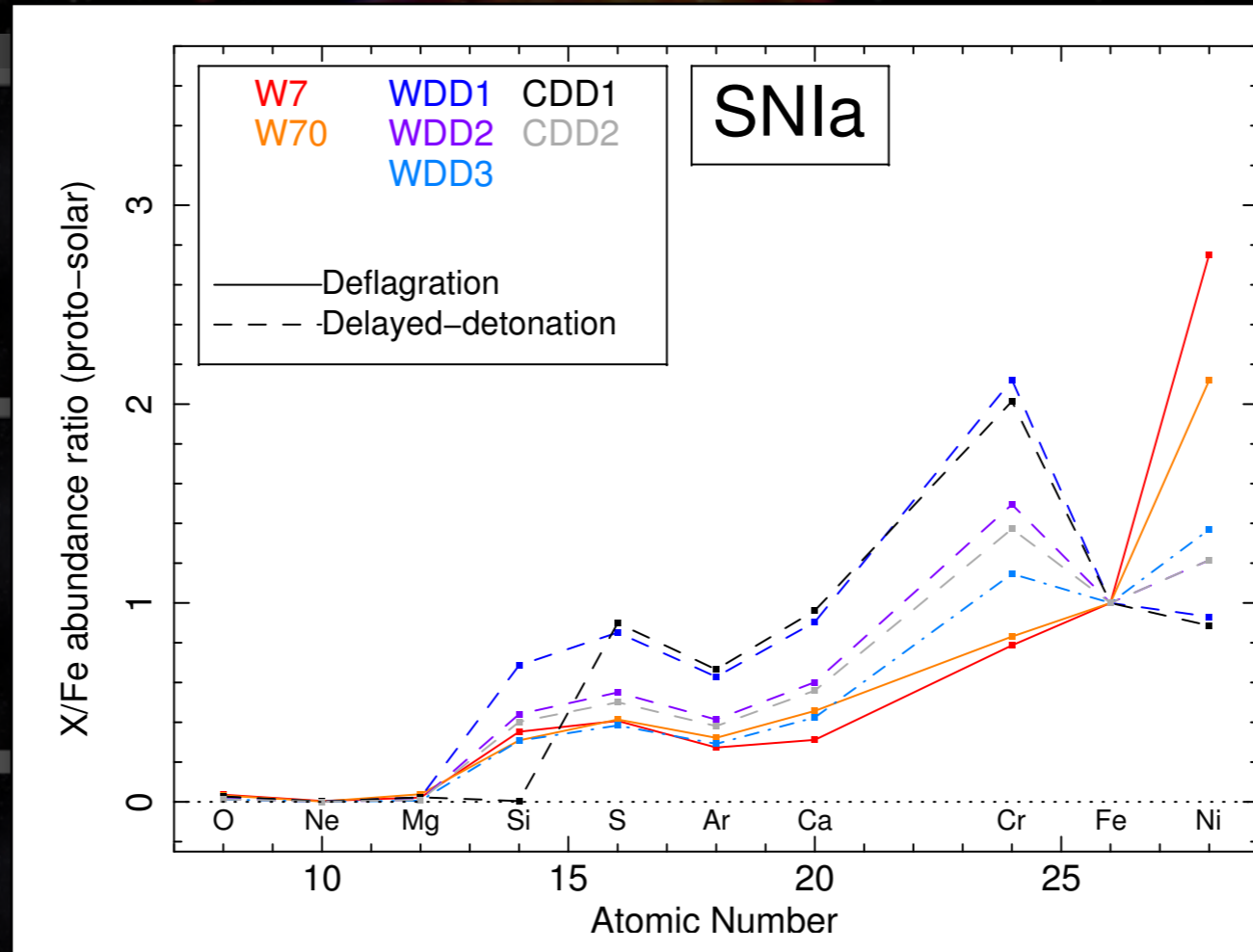
Mennyi? Több mindentől függ:

➔ A robbanás fizikája (*deflagráció vs. késleltetett-detonáció*)

➔ Mi az a szupernóva elődje?

A vegyi elemek eredete

2) Ia típusú (termionukleáris) szupernóvák



→ **Ca**

→ **Fe**

→ **Ni**

lásd Iwamoto et al. (1999)

mindentől függ:

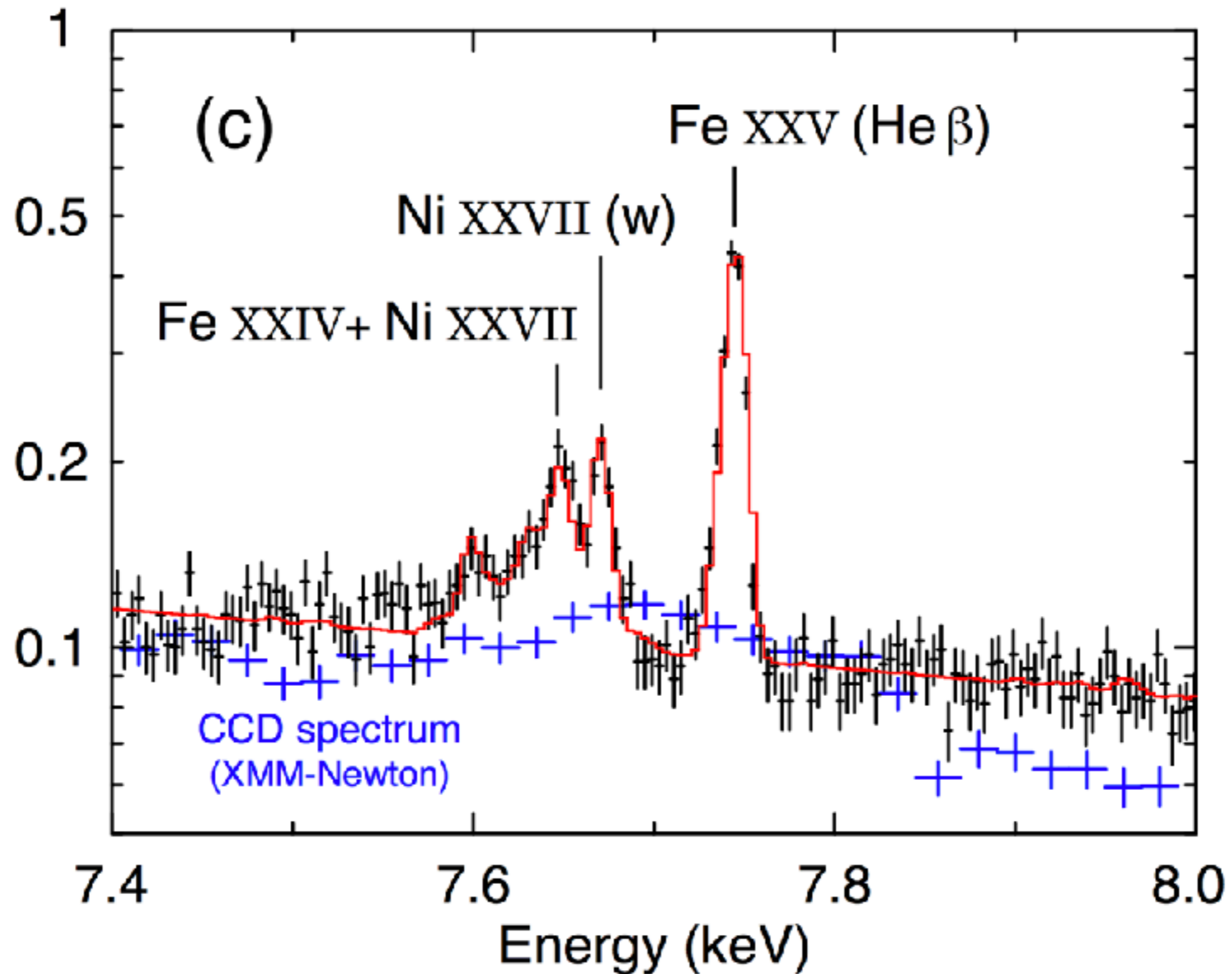
→ A robbanás fizikája (**deflagráció vs. késleltetett-detonáció**)

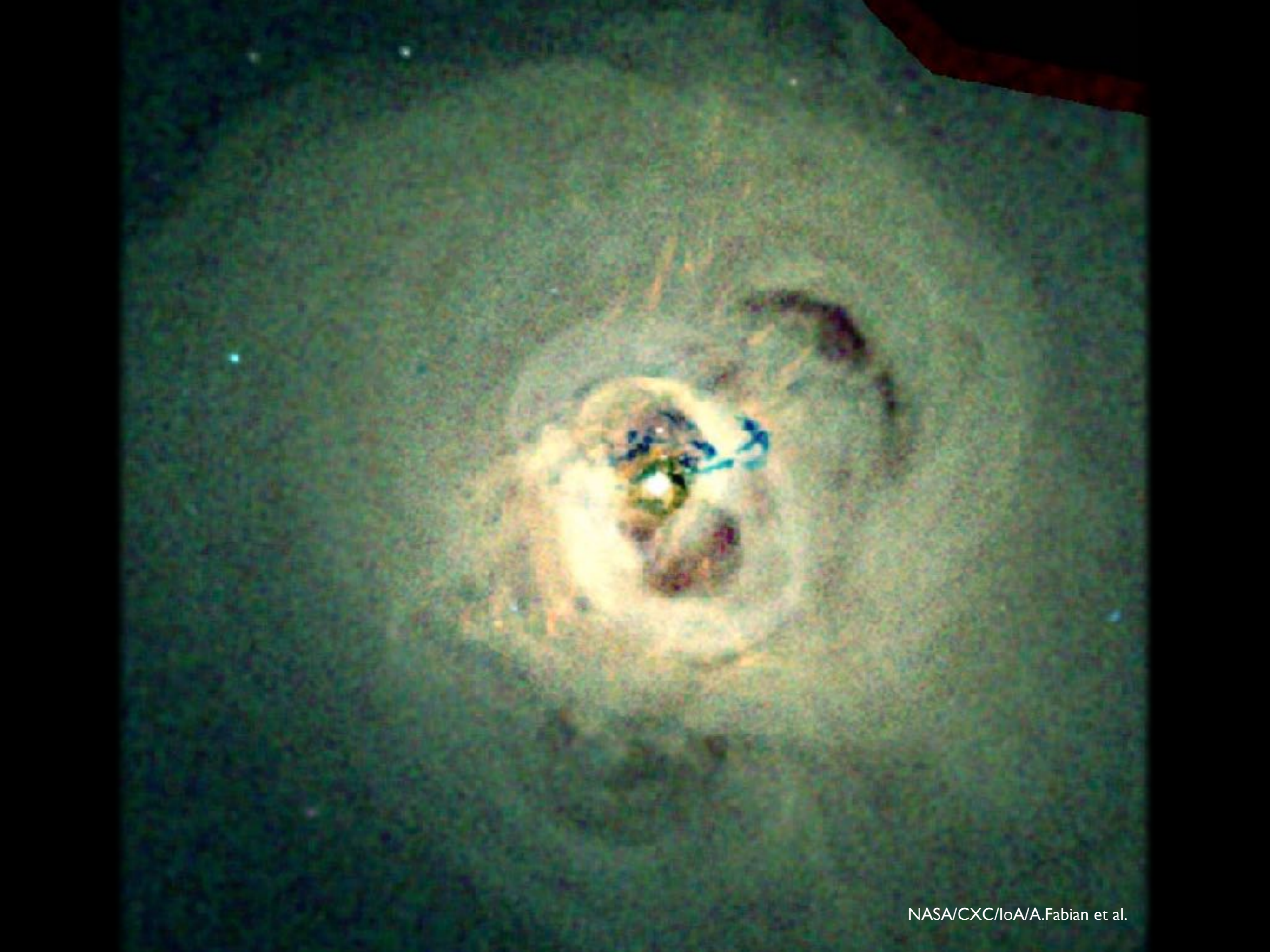
→ Mi az Ia szupernóva elődje?

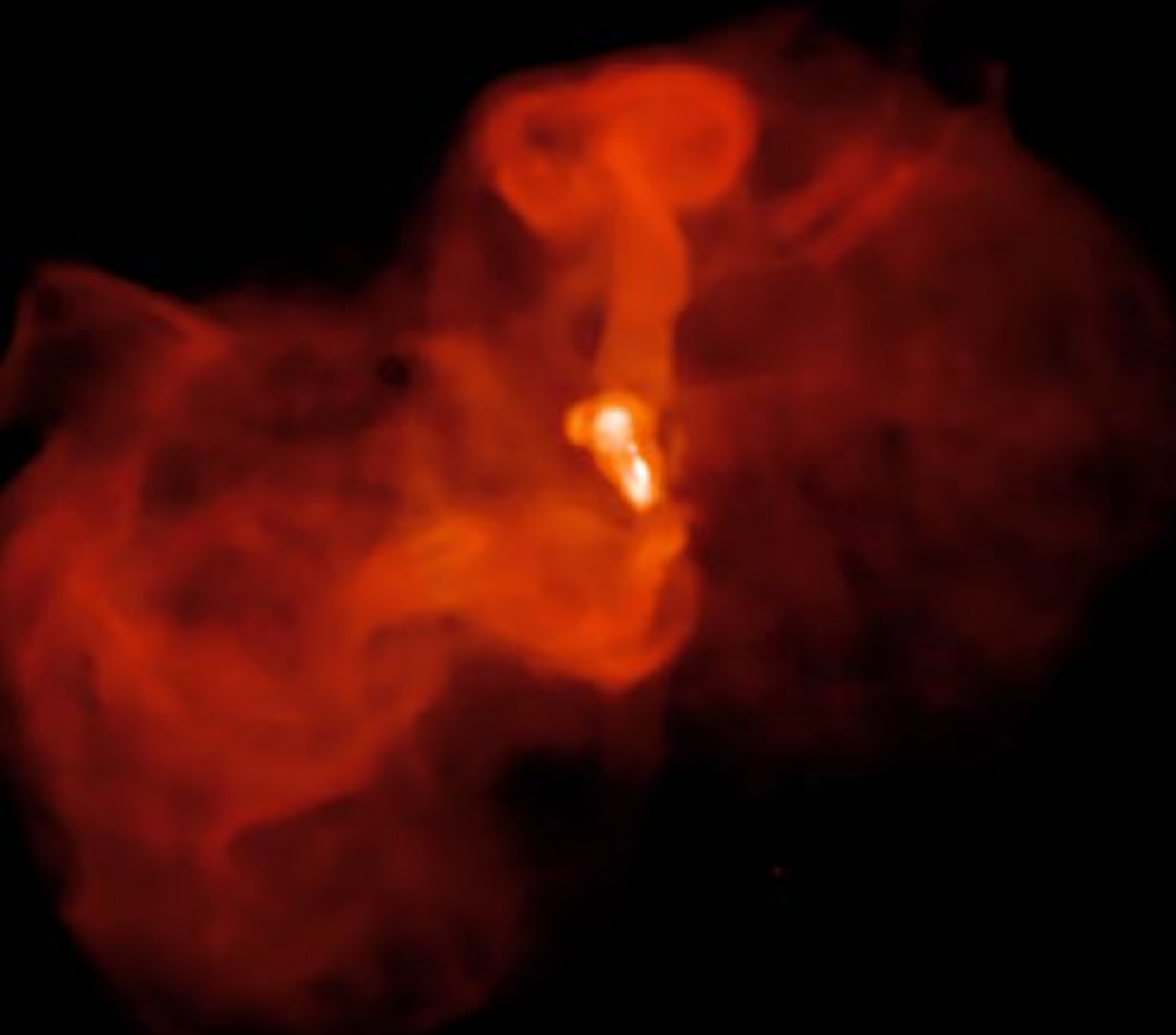
2 fehér törpe?

1 fehér törpe + csillag

A nikkél vonalak felbontása



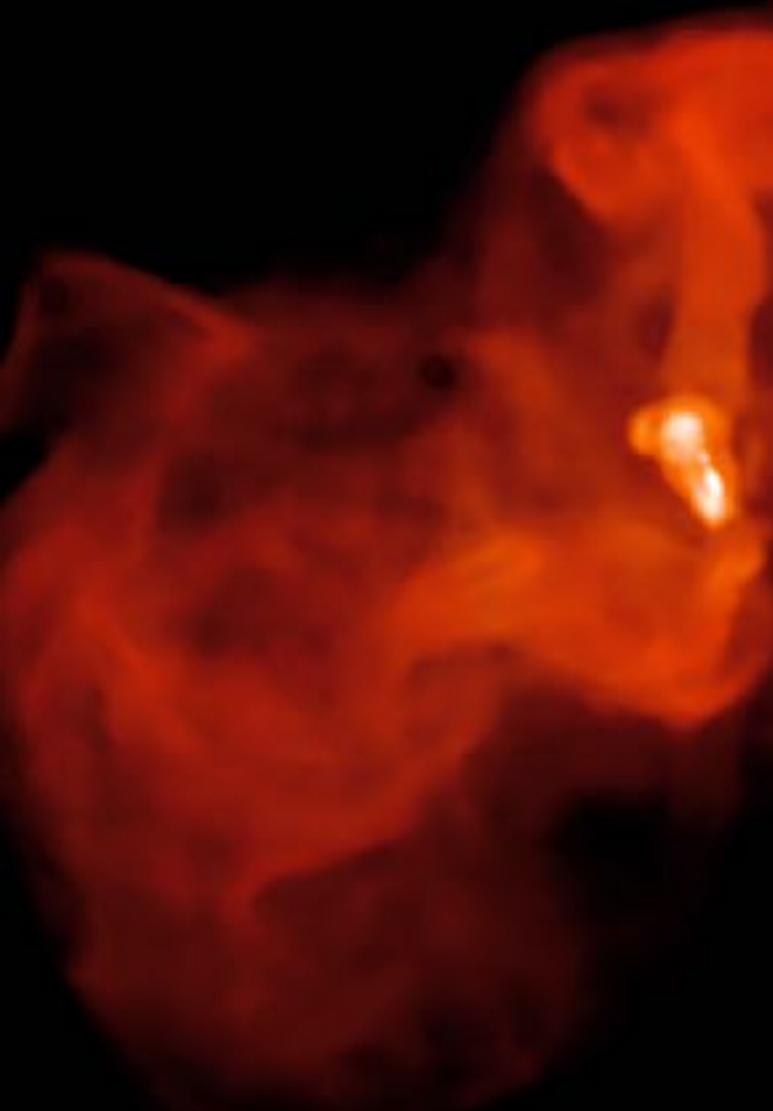




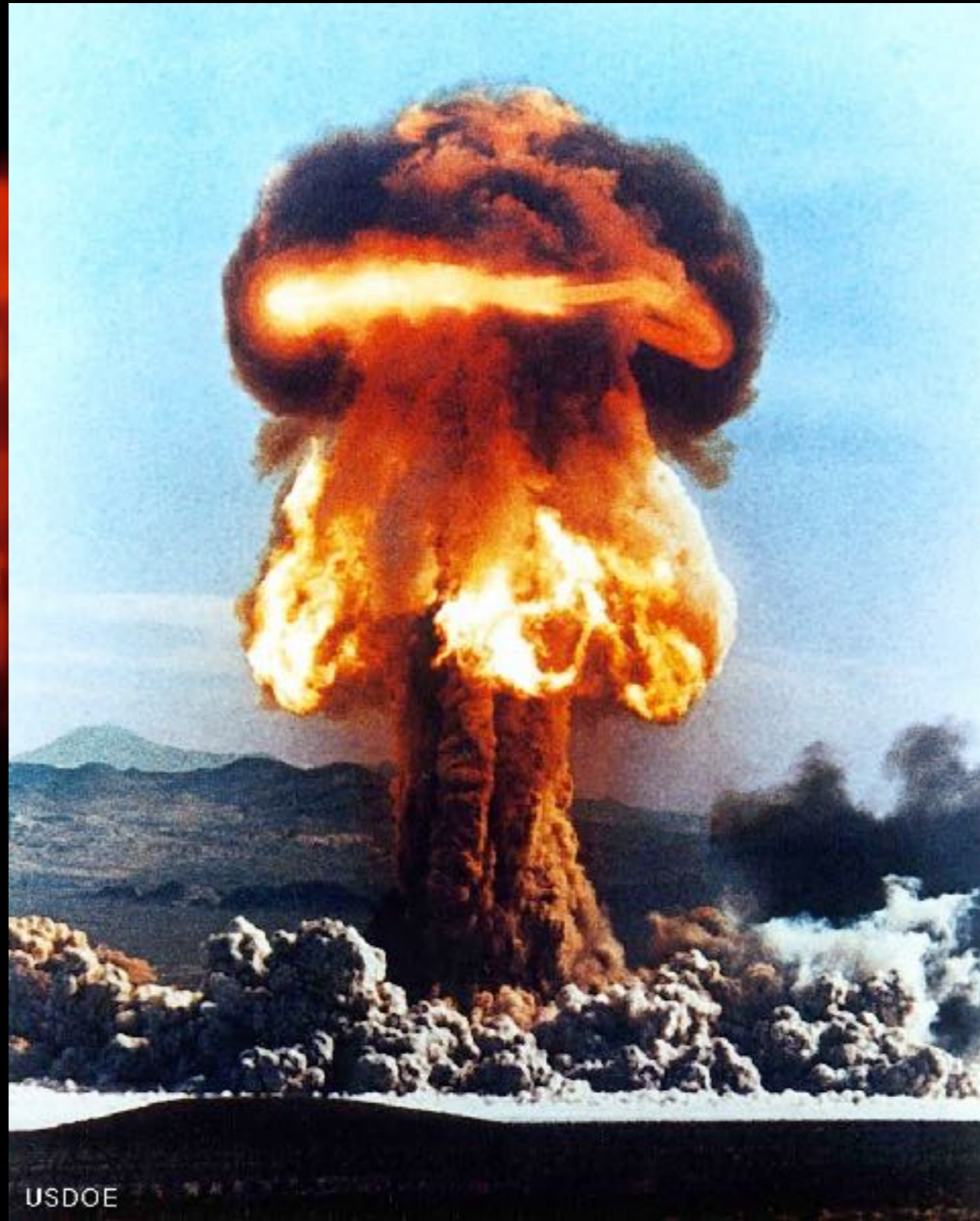
NRAO/AUI/NSF/F. Owen



NASA/CXC/KIPAC/N. Werner, E. Million et al.

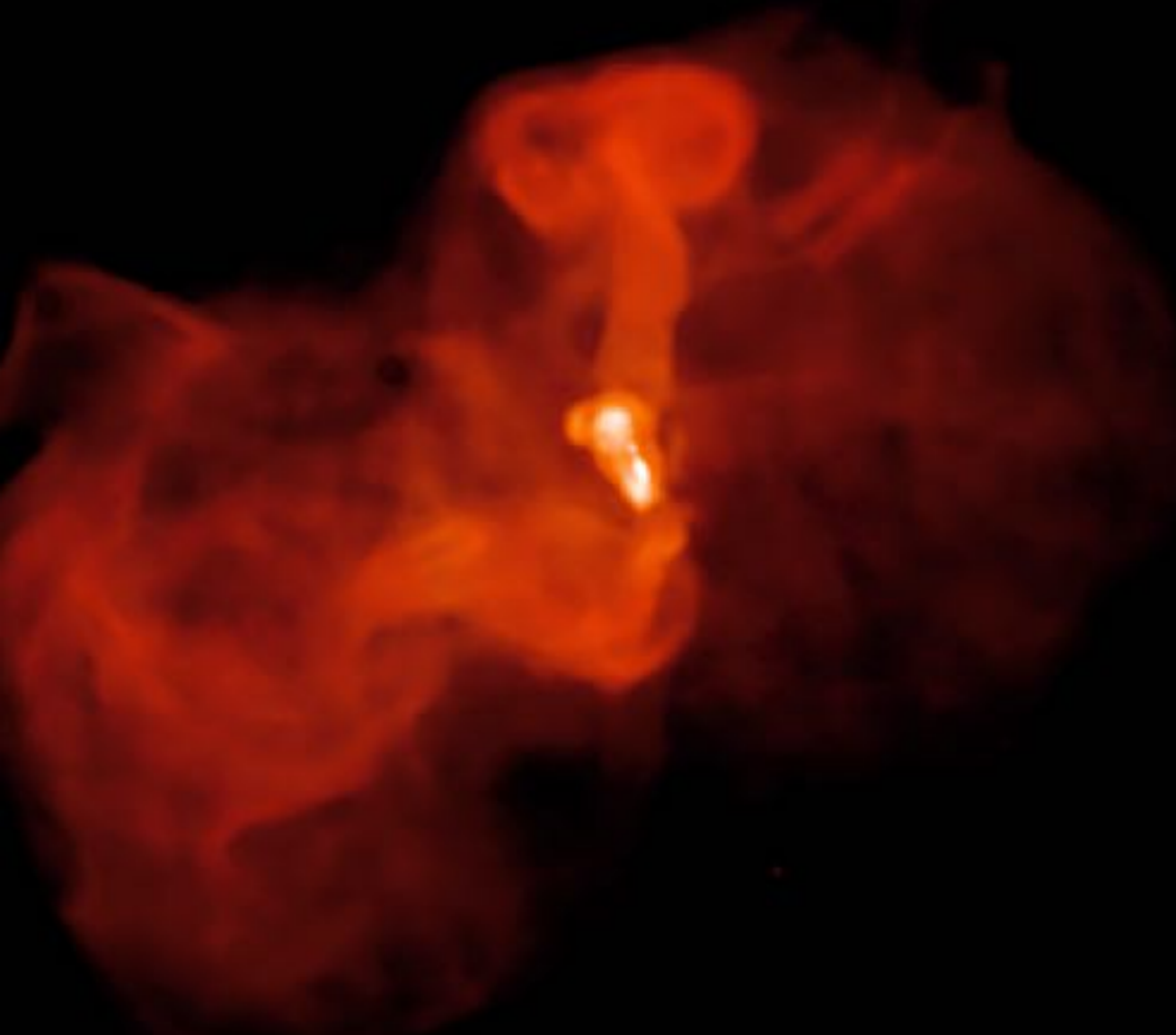


NRAO/AUI/NSF/F. Owen



USDOE





NRAO/AUI/NSF/F. Owen



NASA/CXC/KIPAC/N. Werner, E. Million et al.

Kitörések közel és távol



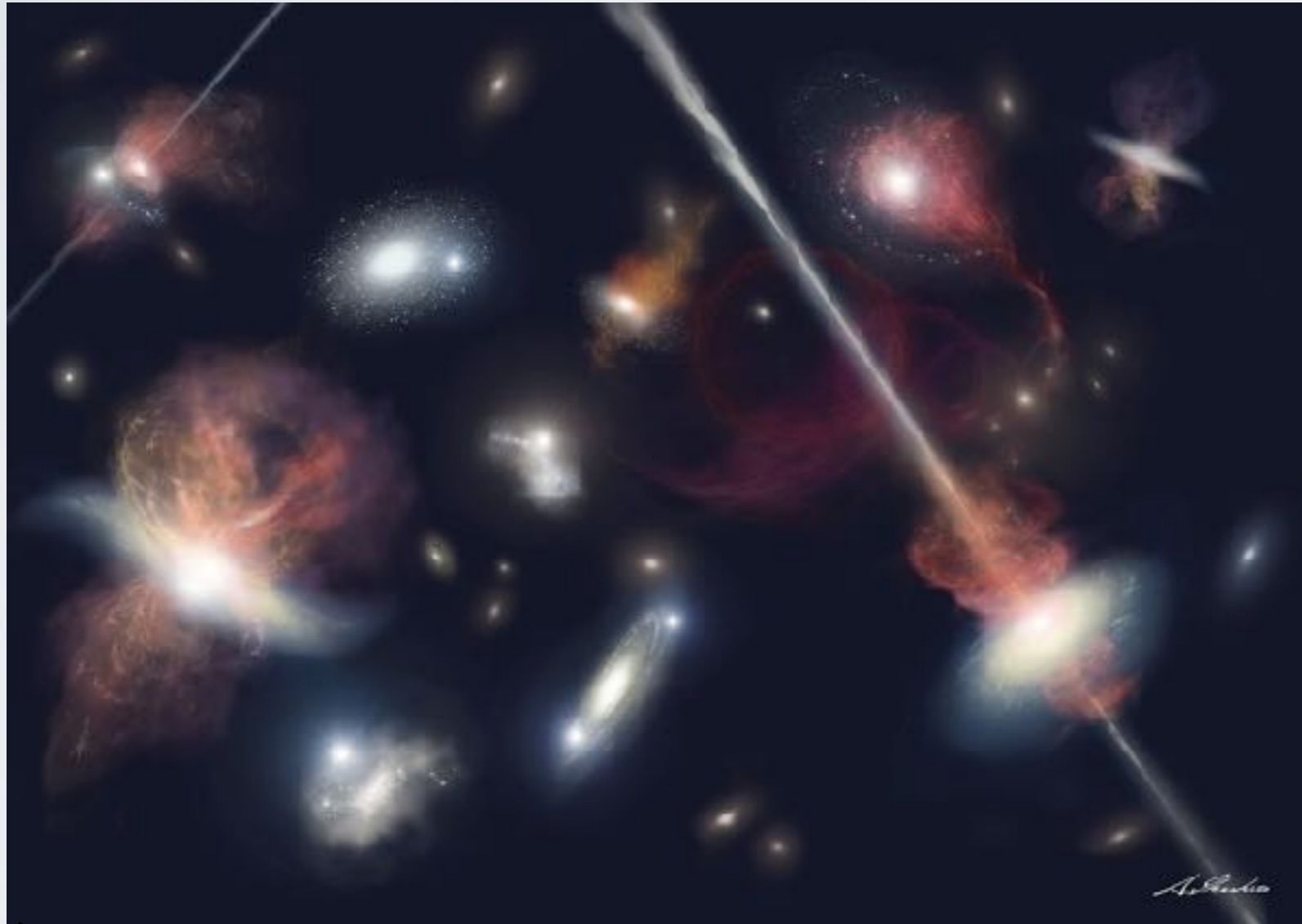
Omar Ragnarsson

Kitörések közel és távol

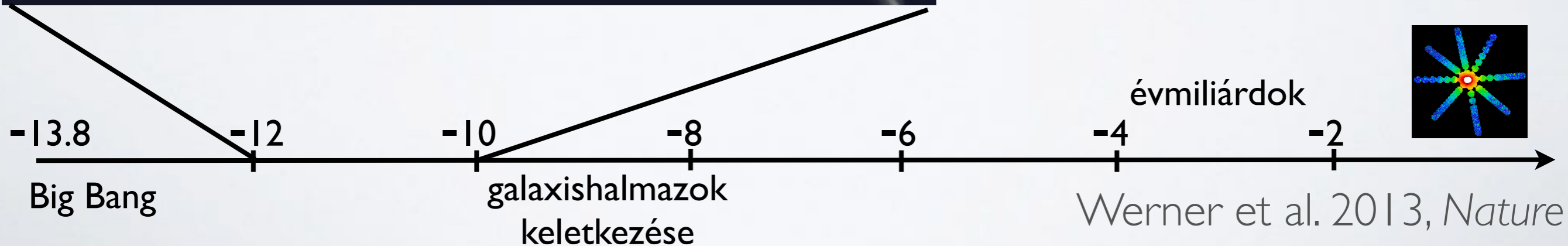


Omar Ragnarsson

A fémek szétszóródása és keveredése

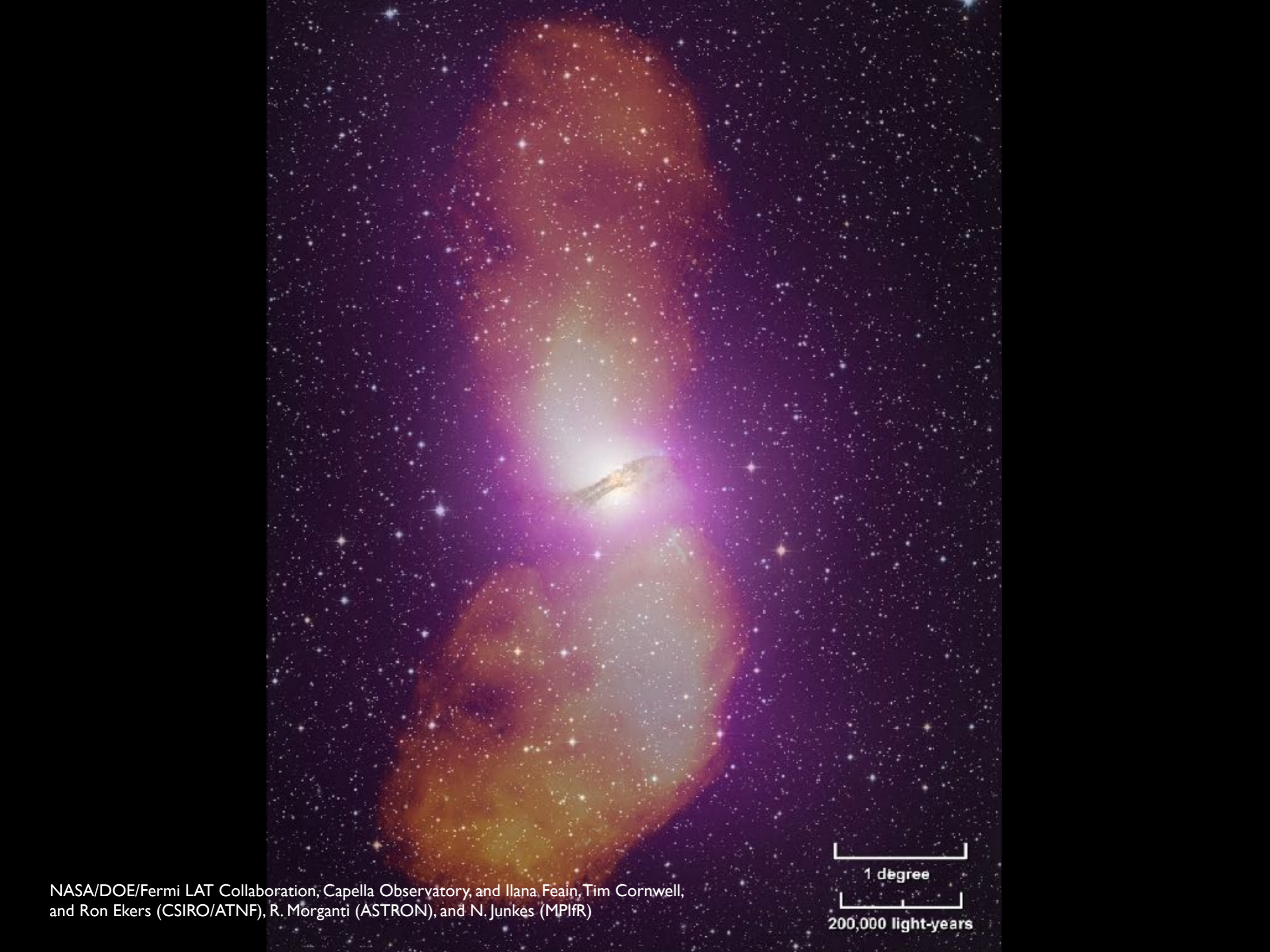


- A galaxisközi gázban található fémek nagy része 10-12 milárd éve keletkezett és szóródott szét világegyetemben.



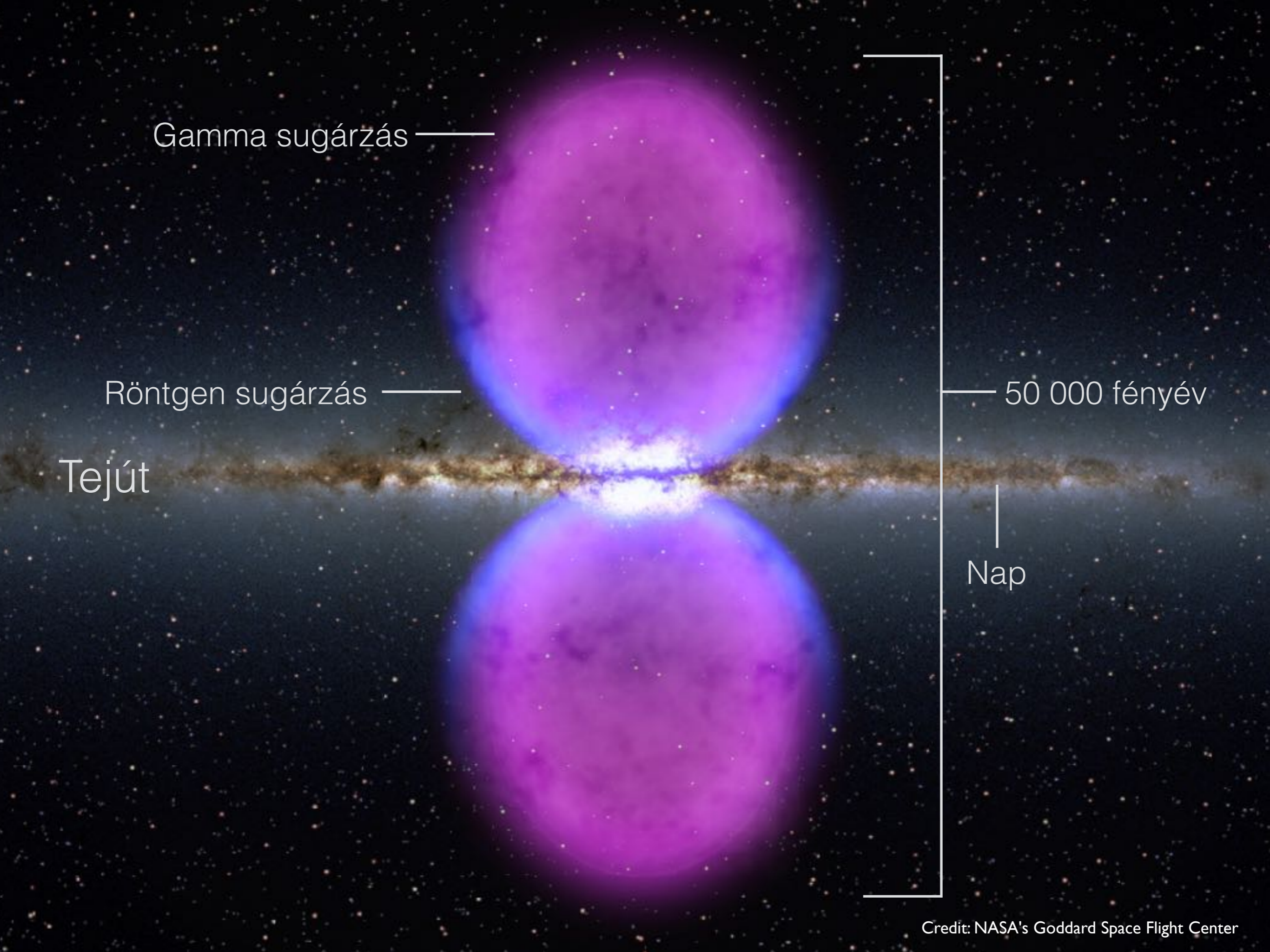






NASA/DOE/Fermi LAT Collaboration, Capella Observatory, and Ilana Feain, Tim Cornwell,
and Ron Ekers (CSIRO/ATNF), R. Morganti (ASTRON), and N. Junkes (MPIfR)

1 degree
200,000 light-years



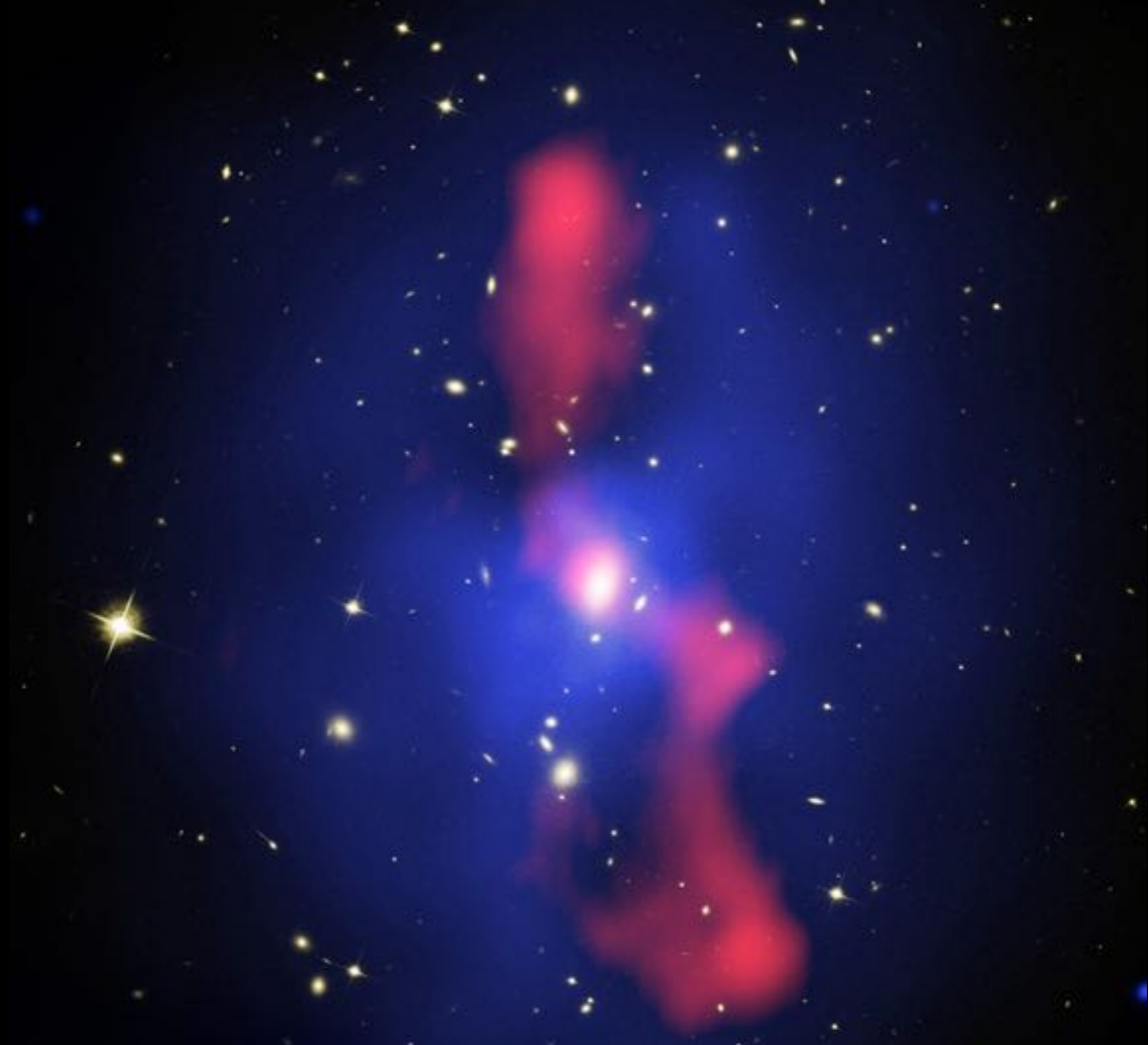
Gamma sugárzás

Röntgen sugárzás

Tejút

50 000 fényév

Nap



X-ray: NASA/CXC/Univ. Waterloo/B. McNamara; Optical: NASA/ESA/STScI/Univ. Waterloo/B. McNamara; Radio: NRAO/Ohio Univ./L. Birzan et al.

600 000 fényév

